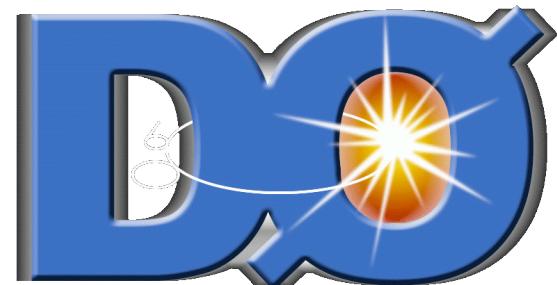
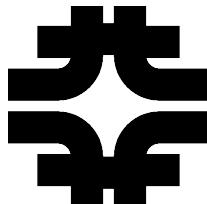


# “Nuts and Bolts” of Combining Tevatron Higgs Search Results



Tom Junk  
Fermilab



- Exchanging Experimental Results
- Bayesian Limits
- $CL_s$  Limits
- Discovery Techniques
- Tevatron Combined Results

<http://tevnphwg.fnal.gov>

<http://www-cdf.fnal.gov/physics/new/hdg/hdg.html>

<http://www-d0.fnal.gov/Run2Physics/WWW/results/higgs.htm>

# Exchanging Experimental Results for Combination

At LEP and at the Tevatron, we exchanged histograms of observed and predicted events.

## Many advantages:

- Crosscheck analyzers' work:  
Signal and background checksum  
Limit/discovery recalculations  
check for “broken” bins
  - $s>0$  when  $b=0$
  - any observation or prediction  $<0$
- Can make control plots
- Can try a great variety of statistical treatments
  - Profile Likelihood, Bayesian,  $CL_s$  and compare each one
- Can make expected limits/LLR distributions without approximations
- Can draw the  $\pm 1\sigma$ ,  $\pm 2\sigma$  bands on expected limits with MC
- Can point to excesses and deficits to explain why limits and p-values are as they are
- Can accommodate new cross sections and branching ratios by scaling
- Can pick and choose signals if more than one expected
  - (e.g., do 4th gen analysis with  $H \rightarrow WW$  without  $WH$ ,  
 $ZH$  and  $VBF$ )
- Pre-binned histograms mean combiners don't have to  
choose binning, reducing mistakes, inconsistencies

## Disadvantages:

- Lots of work/CPU!
- Have to share preliminary histos  
(your competitors may find your mistakes!)

# Exchanging Experimental Results for Combination

## Systematic uncertainties itemized by named source

- Asymmetric Rate errors on each predicted component
- Shape errors supplied as alternate shape histograms
  - Alternate shapes normalized to predicted yield under parameter variation.  
(can have a different rate and shape from the best value)
- practical difficulty: How to estimate  $5\sigma$  systematics?
- Bin-by-bin independent uncertainties (MC statistics)
- Names used to categorize systematic uncertainties in a way easy to understand and check – and to implement correlations

Same name of uncertainty – 100% correlated. Different name: 0% correlated. Can always synthesize arbitrary correlations from mixtures of these.

# Cross Section and Branching Ratio Alignment

Table 4: Higgs boson decay branching fractions computed with HDECAY

SM b.r.'s from

HDECAY 3.15

(April 2010.

Reran the latest

from Jan 2011

and no changes)

We will update to

the b.r.'s in

arXiv:1101.0593v2

(Handbook of LHC

Cross Sections

1. Inclusive Observables)

For Summer 2011

$m_H$ GeV/ $c^2$	$B(H \rightarrow b\bar{b})$ (%)	$B(H \rightarrow c\bar{c})$ (%)	$B(H \rightarrow \tau^+\tau^-)$ (%)	$B(H \rightarrow W^+W^-)$ (%)	$B(H \rightarrow ZZ)$ (%)	$B(H \rightarrow \gamma\gamma)$ (%)
100	80.33	3.542	7.920	1.052	0.1071	0.1505
105	78.57	3.463	7.821	2.307	0.2035	0.1689
110	75.90	3.343	7.622	4.585	0.4160	0.1870
115	71.95	3.169	7.288	8.268	0.8298	0.2029
120	66.49	2.927	6.789	13.64	1.527	0.2148
125	59.48	2.617	6.120	20.78	2.549	0.2204
130	51.18	2.252	5.305	29.43	3.858	0.2182
135	42.15	1.854	4.400	39.10	5.319	0.2077
140	33.04	1.453	3.472	49.16	6.715	0.1897
145	24.45	1.075	2.585	59.15	7.771	0.1653
150	16.71	0.7345	1.778	68.91	8.143	0.1357
155	9.88	0.4341	1.057	78.92	7.297	0.09997
160	3.74	0.1646	0.403	90.48	4.185	0.05365
165	1.29	0.05667	0.140	95.91	2.216	0.02330
170	0.854	0.03753	0.0930	96.39	2.351	0.01598
175	0.663	0.02910	0.0725	95.81	3.204	0.01236
180	0.535	0.02349	0.0589	93.25	5.937	0.01024
185	0.415	0.01823	0.0459	84.50	14.86	0.008128
190	0.340	0.01490	0.0377	78.70	20.77	0.006774
195	0.292	0.01281	0.0326	75.88	23.66	0.005919
200	0.257	0.01128	0.0288	74.26	25.33	0.005285
210	0.207	0.00909	0.0259	72.50	27.16	0.004358
220	0.188	0.00756	0.0235	71.60	28.11	0.003677
230	0.146	0.00640	0.0197	71.05	28.70	0.003147
240	0.125	0.00550	0.0168	70.66	29.11	0.002720
250	0.109	0.00477	0.0145	70.36	29.43	0.002370
260	0.0953	0.00418	0.0127	70.12	29.70	0.002078
270	0.0843	0.00370	0.00999	69.91	29.92	0.001833
280	0.0750	0.00329	0.00895	69.72	30.12	0.001625
290	0.0672	0.00295	0.00807	69.56	30.29	0.001448
300	0.0605	0.00266	0.00732	69.40	30.45	0.001297

All channels must use consistent predictions

TABLE IV: Higgs boson decay branching fractions from the LHC Cross Section Handbook [12]

$m_H$ GeV/ $c^2$	$B(H \rightarrow b\bar{b})$ (%)	$B(H \rightarrow c\bar{c})$ (%)	$B(H \rightarrow \tau^+\tau^-)$ (%)	$B(H \rightarrow W^+W^-)$ (%)	$B(H \rightarrow ZZ)$ (%)	$B(H \rightarrow \gamma\gamma)$ (%)
90	81.2	3.78	8.41	0.209	0.0421	0.123
95	80.4	3.73	8.41	0.472	0.0672	0.140
100	79.1	3.68	8.36	1.11	0.113	0.159
105	77.3	3.59	8.25	2.43	0.215	0.178
110	74.5	3.46	8.03	4.82	0.439	0.197
115	70.5	3.27	7.65	8.67	0.873	0.213
120	64.9	3.01	7.11	14.3	1.60	0.225
125	57.8	2.68	6.37	21.6	2.67	0.230
130	49.4	2.29	5.49	30.5	4.02	0.226
135	40.4	1.87	4.52	40.3	5.51	0.214
140	31.4	1.46	3.54	50.4	6.92	0.194
145	23.1	1.07	2.62	60.3	7.96	0.168
150	15.7	0.725	1.79	69.9	8.28	0.137
155	9.18	0.425	1.06	79.6	7.36	0.100
160	3.44	0.159	0.397	90.9	4.16	0.0533
165	1.19	0.0549	0.138	96.0	2.22	0.0230
170	0.787	0.0364	0.0920	96.5	2.36	0.0158
175	0.612	0.0283	0.0719	95.8	3.23	0.0123
180	0.497	0.0230	0.0587	93.2	6.02	0.0102
185	0.385	0.0178	0.0457	84.4	15.0	0.00809
190	0.315	0.0146	0.0376	78.6	20.9	0.00674
195	0.270	0.0125	0.0324	75.7	23.9	0.00589
200	0.238	0.0110	0.0287	74.1	25.6	0.00526
210	0.192	0.00889	0.0234	72.3	27.4	0.00434
220	0.160	0.00740	0.0196	71.4	28.4	0.00367
230	0.136	0.00627	0.0168	70.8	28.9	0.00314
240	0.117	0.00539	0.0145	70.4	29.4	0.00272
250	0.101	0.00468	0.0127	70.1	29.7	0.00237
260	0.0889	0.00411	0.0112	69.9	29.9	0.00208
270	0.0786	0.00363	0.0100	69.7	30.2	0.00184
280	0.0700	0.00323	0.00898	69.5	30.4	0.00163
290	0.0627	0.00290	0.00809	69.3	30.5	0.00145
300	0.0565	0.00261	0.00733	69.2	30.7	0.00130

# Uncertainties in Higgs Boson Decay Branching Ratios

Baglio and Djouadi,  
arXiv:1012.0530

plus private communication  
to get the entire table

$m_b$  – raising  $m_b$  increases  
the  $b\bar{b}$  decay rate

$m_c$  – same thing for charm

$\alpha_s$  – increasing this raises  
 $\text{Br}(H \rightarrow gg)$ .

Sum  $\text{Br.}=1$  means  
 $\text{Br}(H \rightarrow WW)$  is impacted  
accordingly.

“Up” uncertainty corresponds  
to a positive shift in the  
named parameter.

$m_H$ (GeV/c <sup>2</sup> )	$\delta \text{Br}(H \rightarrow W^+W^-)$ ( $\alpha_s$ ) (%)	$\delta \text{Br}(H \rightarrow W^+W^-)$ ( $m_b$ ) (%)	$\delta \text{Br}(H \rightarrow W^+W^-)$ ( $m_c$ ) (%)
100	-1.57 +1.67	-7.61 +2.78	-0.74 +0.93
105	-1.67 +1.75	-7.55 +2.66	-0.72 +0.83
110	-1.60 +1.68	-7.29 +2.60	-0.68 +0.80
115	-1.52 +1.57	-6.91 +2.44	-0.65 +0.75
120	-1.43 +1.43	-6.38 +2.24	-0.61 +0.68
125	-1.26 +1.26	-5.75 +1.98	-0.54 +0.58
130	-1.02 +1.09	-4.89 +1.70	-0.45 +0.51
135	-0.83 +0.88	-4.04 +1.39	-0.36 +0.44
140	-0.66 +0.66	-3.17 +1.06	-0.29 +0.33
145	-0.46 +0.49	-2.33 +0.79	-0.20 +0.25
150	-0.33 +0.31	-1.59 +0.51	-0.14 +0.16
155	-0.18 +0.19	-0.94 +0.31	-0.08 +0.10
160	-0.07 +0.07	-0.35 +0.11	-0.03 +0.03
165	-0.02 +0.02	-0.12 +0.04	-0.01 +0.01
170	-0.01 +0.01	-0.08 +0.02	-0.01 +0.01
175	-0.01 +0.01	-0.06 +0.02	+0.00 +0.01
180	-0.01 +0.01	-0.04 +0.02	+0.00 +0.01
185	-0.01 +0.01	-0.04 +0.01	+0.00 0.00
190	+0.00 +0.01	-0.03 +0.01	+0.00 0.00
195	+0.00 +0.01	-0.03 +0.01	+0.00 0.00
200	+0.00 +0.01	-0.01 +0.01	+0.00 0.00
210	0.00 +0.00	-0.01 -0.01	+0.00 +0.00
220	0.00 +0.00	0.00 -0.01	0.00 +0.00
230	0.00 -0.01	+0.01 -0.01	0.00 +0.00
240	0.00 +0.00	0.00 +0.00	0.00 +0.00
250	0.00 +0.00	+0.00 +0.01	0.00 0.00
260	0.00 +0.00	0.00 -0.01	0.00 +0.00
270	0.00 +0.00	0.00 -0.01	0.00 +0.00
280	0.00 +0.00	0.00 +0.00	0.00 +0.00
290	0.00 +0.00	+0.00 +0.01	0.00 0.00
300	0.00 +0.00	-0.01 0.00	+0.00 0.00

# The Same Uncertainties for $\text{Br}(\text{H} \rightarrow \text{bb})$

Relative signs preserved  
with respect to the  $\text{Br}(\text{H} \rightarrow \text{WW})$   
table on the previous page.

These only became available  
after our most recent low-mass  
limits were set. To be  
included next time. The WW  
uncertainties went into our  
high-mass limits.

Also have these tables for  
 $\text{Br}(\text{H} \rightarrow \text{cc})$ ,  $\text{Br}(\text{H} \rightarrow \tau\tau)$

Baglio and  
Djouadi

$m_H$ (GeV/c <sup>2</sup> )	$\delta\text{Br}(H \rightarrow bb)$ ( $\alpha_s$ ) (%)	$\delta\text{Br}(H \rightarrow bb)$ ( $m_b$ ) (%)	$\delta\text{Br}(H \rightarrow bb)(m_c)$ (%)
100	+0.34 -0.35	+1.99 -0.70	-0.74 +0.88
105	+0.40 -0.40	+2.20 -0.77	-0.72 +0.85
110	+0.47 -0.49	+2.47 -0.88	-0.71 +0.81
115	+0.58 -0.61	+2.90 -1.02	-0.66 +0.78
120	+0.72 -0.75	+3.47 -1.21	-0.61 +0.71
125	+0.90 -0.94	+4.23 -1.45	-0.54 +0.64
130	+1.12 -1.14	+5.12 -1.75	-0.47 +0.55
135	+1.32 -1.37	+6.05 -2.07	-0.40 +0.42
140	+1.54 -1.54	+7.04 -2.35	-0.29 +0.35
145	+1.71 -1.75	+7.94 -2.68	-0.26 +0.26
150	+1.88 -1.94	+8.74 -2.91	-0.19 +0.13
155	+2.03 -2.06	+9.51 -3.08	-0.10 +0.11
160	+2.17 -2.17	+10.17 -3.26	-0.03 +0.06
165	+2.23 -2.23	+10.45 -3.34	+0.00 +0.09
170	+2.24 -2.25	+10.45 -3.36	-0.03 +0.03
175	+2.25 -2.25	+10.48 -3.37	-0.03 +0.02
180	+2.25 -2.27	+10.48 -3.37	-0.02 +0.02
185	+2.26 -2.26	+10.50 -3.36	-0.03 +0.03
190	+2.28 -2.28	+10.49 -3.39	-0.03 +0.03
195	+2.31 -2.28	+10.50 -3.38	+0.00 +0.04
200	+2.28 -2.28	+10.51 -3.36	+0.00 0.00
210	+2.30 -2.30	+10.52 -3.37	+0.00 0.00
220	+2.32 -2.32	+10.55 -3.34	+0.00 0.00
230	+2.28 -2.35	+10.48 -3.42	+0.00 0.00
240	+2.30 -2.39	+10.52 -3.36	+0.00 0.00
250	+2.36 -2.35	+10.51 -3.38	-0.02 +0.02
260	+2.36 -2.37	+10.53 -3.38	-0.02 +0.02
270	+2.38 -2.37	+10.54 -3.38	-0.01 +0.03
280	+2.38 -2.38	+10.54 -3.37	-0.01 +0.03
290	+2.39 -2.39	+10.53 -3.38	-0.02 +0.02
300	+2.40 -2.40	+10.53 -3.37	-0.02 +0.02

# Cross Section and Branching Ratio Alignment

$gg \rightarrow H$  production cross sections and uncertainties.

From M. Grazzini and D. de Florian  
Similar calc. from Anastasiou,  
Boughezal, and Petriello

Updated:  $m_t = 173.1$  GeV,  
full  $m_t$  dependence.

In the main channels, we ignore the scale and PDF uncertainties in this table as they use  $n_{jet}$  categories, and the jet-by-jet uncertainties are larger.

33% relative uncertainty in the 2-jet bin smaller than the one recommended by Anastasiou, Dissertori, Grazzini, and Webber, because of a new NLO calculation of H+2j by Campbell, Ellis, and Williams

Table 3: The SM NNLO+NNLL production cross section for  $gg \rightarrow H$  and the associated uncertainties. The PDF+ $\alpha_s$  uncertainties are the full PDF4LHC prescription.

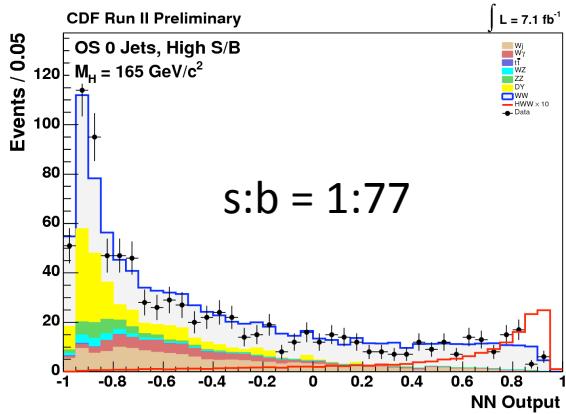
$m_H$ (GeV/ $c^2$ )	$\sigma_{gg \rightarrow H}$ (fb)	scale (up) (%)	scale (down) (%)	PDF+ $\alpha_s$ (up) (%)	PDF+ $\alpha_s$ (down) (%)
90	2442.3	7.39	-8.17	8.70	-11.82
95	2101.1	7.33	-8.02	8.94	-11.84
100	1821.8	7.31	-7.88	9.17	-11.85
105	1584.7	7.28	-7.75	9.41	-11.86
110	1385.0	7.24	-7.64	9.64	-11.87
115	1215.9	7.20	-7.53	9.80	-11.94
120	1072.3	7.19	-7.43	9.96	-12.01
125	949.3	7.13	-7.32	10.07	-12.17
130	842.9	7.18	-7.24	10.18	-12.34
135	750.8	7.18	-7.15	10.39	-12.41
140	670.6	7.17	-7.07	10.59	-12.49
145	600.6	7.16	-7.00	10.80	-12.59
150	539.1	7.16	-6.93	11.01	-12.69
155	484.0	7.15	-6.86	11.23	-12.75
160	432.3	7.16	-6.79	11.45	-12.80
165	383.7	7.18	-6.73	11.62	-12.85
170	344.0	7.21	-6.66	11.79	-12.90
175	309.7	7.23	-6.60	11.90	-12.94
180	279.2	7.23	-6.55	12.00	-12.98
185	252.1	7.24	-6.49	12.10	-13.03
190	228.0	7.25	-6.44	12.20	-13.07
195	207.2	7.28	-6.39	12.26	-13.08
200	189.1	7.30	-6.37	12.31	-13.10
210	158.9	7.37	-6.37	12.26	-12.94
220	134.5	7.43	-6.37	12.53	-13.17
230	114.7	7.48	-6.36	12.61	-13.21
240	98.4	7.55	-6.37	12.64	-13.23
250	85.0	7.65	-6.39	12.81	-13.38
260	73.8	7.76	-6.42	13.02	-13.58
270	64.5	7.83	-6.43	13.22	-13.78
280	56.7	7.93	-6.44	13.75	-14.33
290	50.1	8.06	-6.48	14.27	-14.86
300	44.7	8.23	-6.53	14.90	-15.51
320	36.4	8.44	-6.60	16.05	-16.67
340	31.1	8.74	-6.76	17.62	-18.32
360	29.5	9.05	-6.94	19.93	-20.71
380	25.0	9.35	-7.23	22.34	-23.23
400	19.8	9.66	-7.47	24.81	-25.85

# gg $\rightarrow$ H: No Jet Veto, but there is Classification by Jet Count

## WW+0 jets

CDF Run II Preliminary $\int \mathcal{L} = 7.1 \text{ fb}^{-1}$		
	$M_H = 165 \text{ GeV}/c^2$	
$t\bar{t}$	2.60	$\pm$ 0.79
DY	271	$\pm$ 74
WW	673	$\pm$ 66
WZ	30.3	$\pm$ 4.6
ZZ	45.4	$\pm$ 6.4
W+jets	273	$\pm$ 64
$W\gamma$	169	$\pm$ 25
<b>Total Background</b>	1470	$\pm$ 140
$gg \rightarrow H$	19.7	$\pm$ 3.0
$WH$	0.487	$\pm$ 0.083
$ZH$	0.493	$\pm$ 0.070
$VBF$	0.166	$\pm$ 0.033
<b>Total Signal</b>	20.8	$\pm$ 3.0
<b>Data</b>	1465	

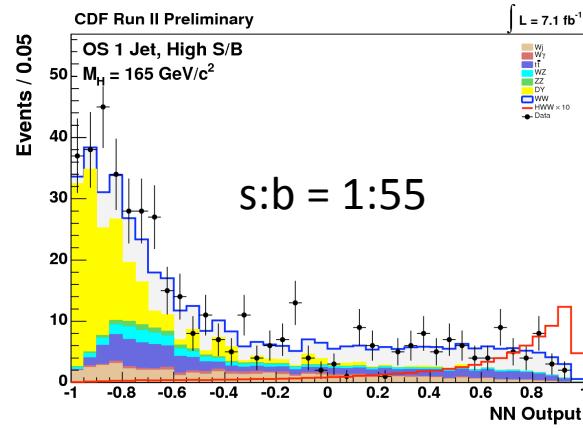
AllSB-0JOS



## WW+1 jet

CDF Run II Preliminary $\int \mathcal{L} = 7.1 \text{ fb}^{-1}$		
	$M_H = 165 \text{ GeV}/c^2$	
$t\bar{t}$	68	$\pm$ 14
DY	264	$\pm$ 59
WW	181	$\pm$ 23
WZ	30.1	$\pm$ 4.2
ZZ	12.3	$\pm$ 1.8
W+jets	94	$\pm$ 24
$W\gamma$	27.5	$\pm$ 4.7
<b>Total Background</b>	676	$\pm$ 84
$gg \rightarrow H$	9.7	$\pm$ 3.0
$WH$	1.34	$\pm$ 0.21
$ZH$	0.521	$\pm$ 0.079
$VBF$	0.88	$\pm$ 0.15
<b>Total Signal</b>	12.4	$\pm$ 3.1
<b>Data</b>	637	

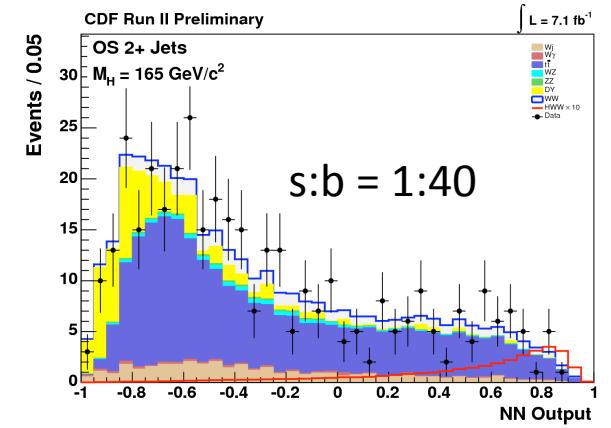
AllSB-1JOS



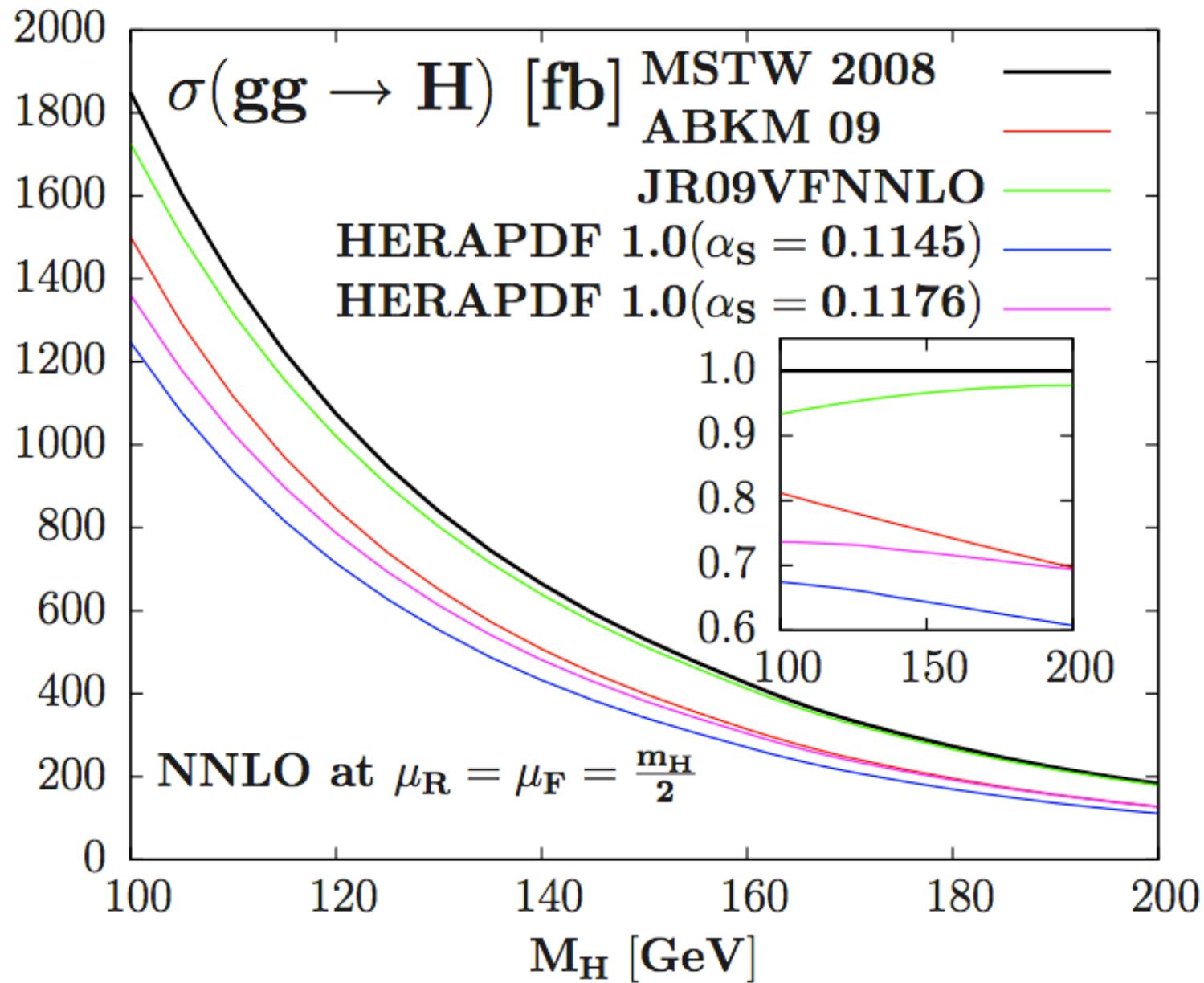
## WW+2 or more jets

CDF Run II Preliminary $\int \mathcal{L} = 7.1 \text{ fb}^{-1}$		
	$M_H = 165 \text{ GeV}/c^2$	
$t\bar{t}$	194	$\pm$ 32
DY	96	$\pm$ 38
WW	40.2	$\pm$ 8.0
WZ	8.2	$\pm$ 1.5
ZZ	3.71	$\pm$ 0.68
W+jets	34.7	$\pm$ 9.7
$W\gamma$	4.9	$\pm$ 1.3
<b>Total Background</b>	382	$\pm$ 61
$gg \rightarrow H$	3.4	$\pm$ 2.6
$WH$	2.97	$\pm$ 0.41
$ZH$	1.52	$\pm$ 0.21
$VBF$	1.63	$\pm$ 0.27
<b>Total Signal</b>	9.5	$\pm$ 2.8
<b>Data</b>	369	

AllSB-2JOS

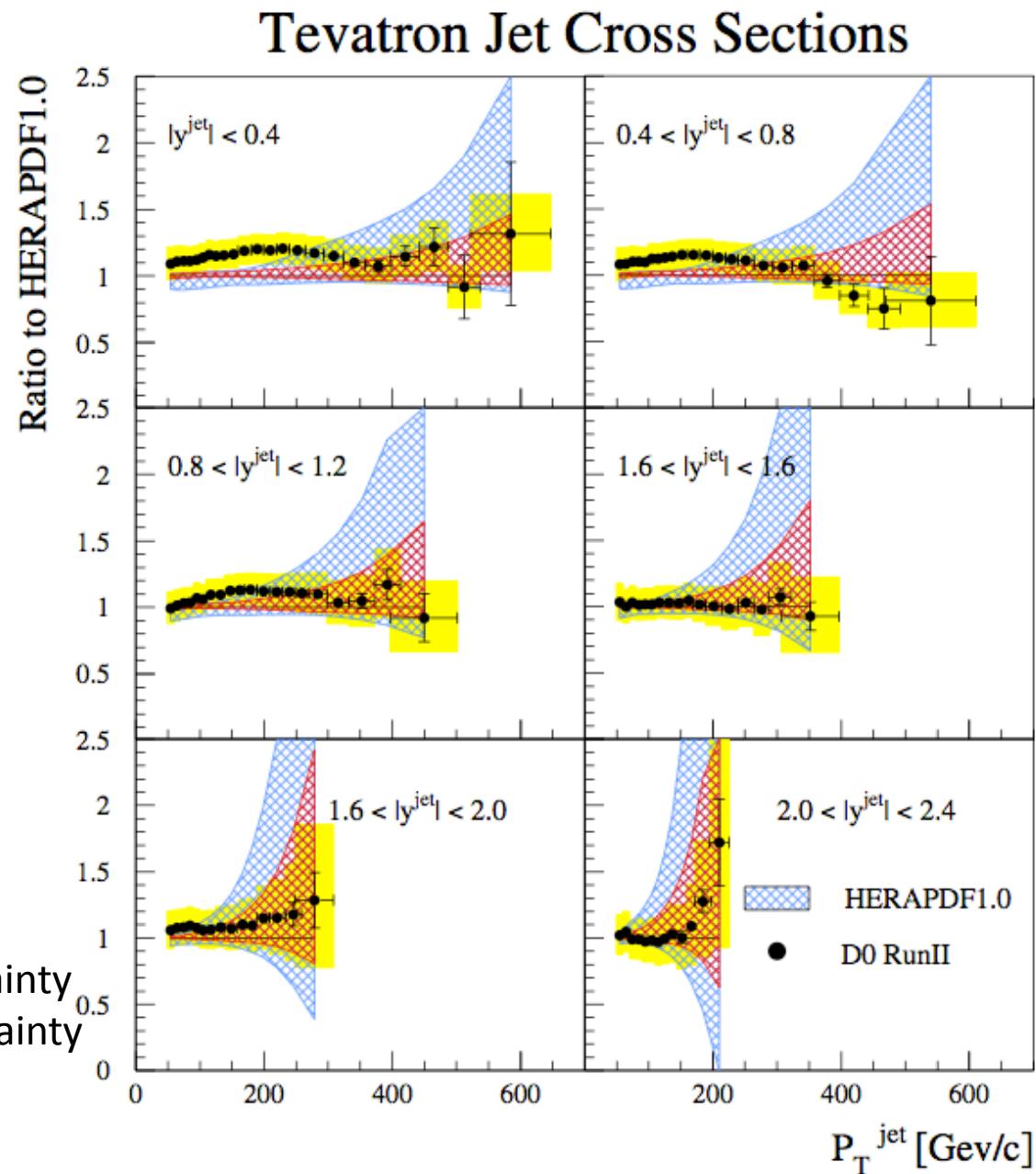


Baglio and Djouadi, arXiv:1101.1832

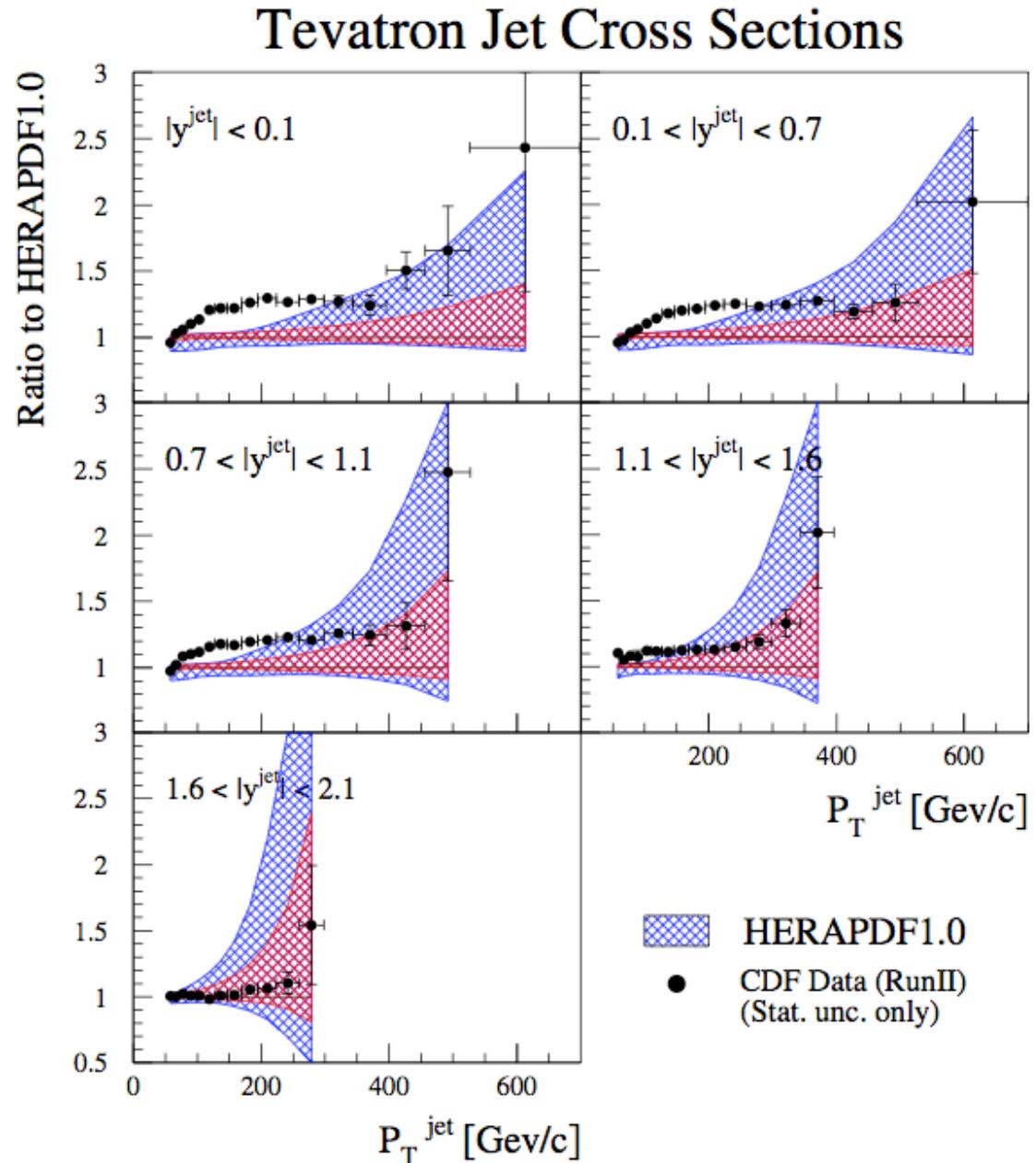


# HERAPDF1.0 Predictions of Tevatron (D0) High- $E_T$ jet data

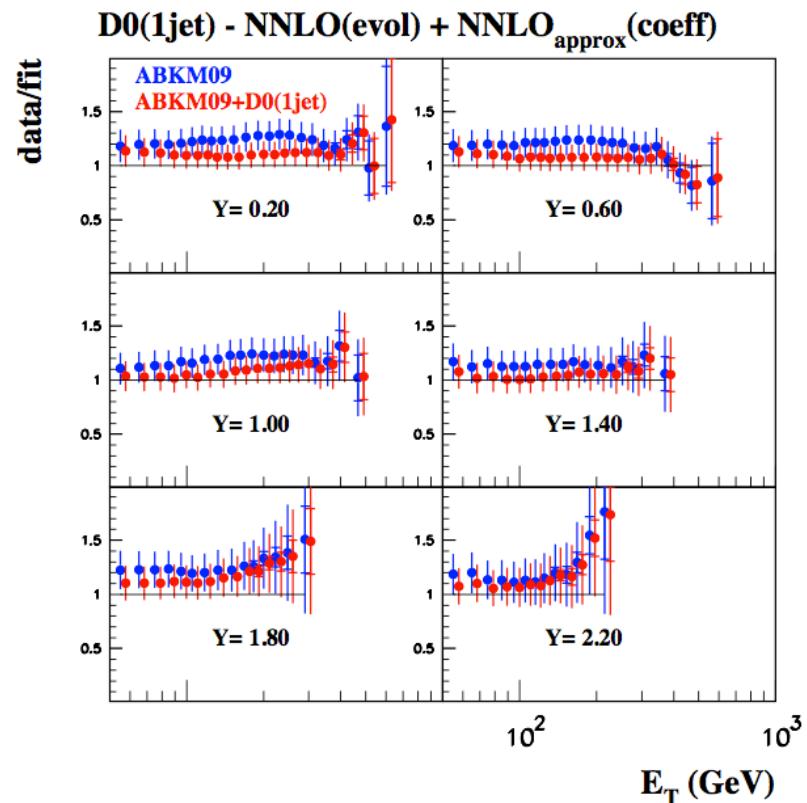
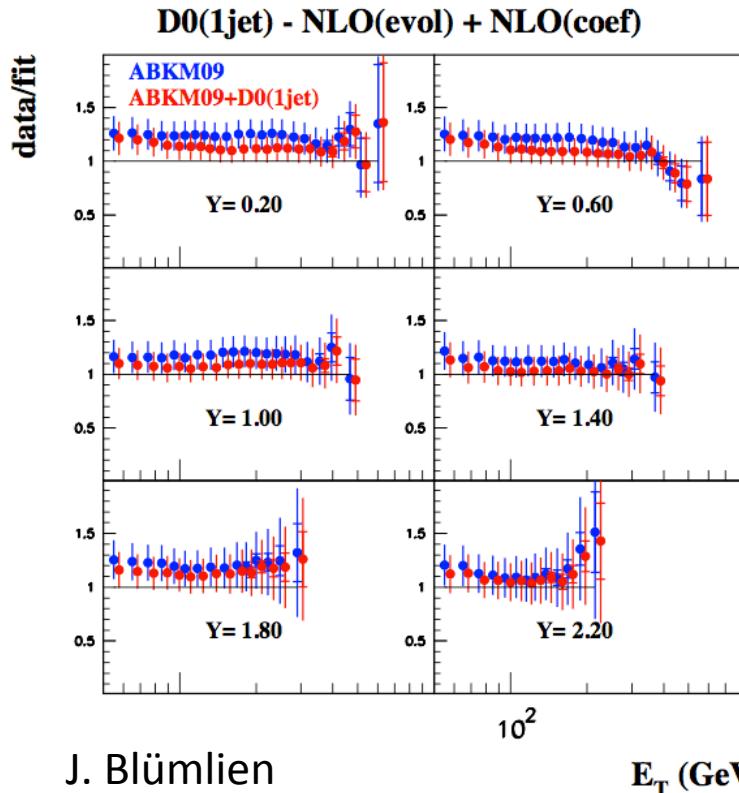
Blue: Total uncertainty,  
Red: Experimental Uncertainty  
Yellow: D0 jet data uncertainty



# HERAPDF1.0 Predictions of Tevatron (CDF) High- $E_T$ jet data



# ABKM09 fit including D0 Run II Jet Data

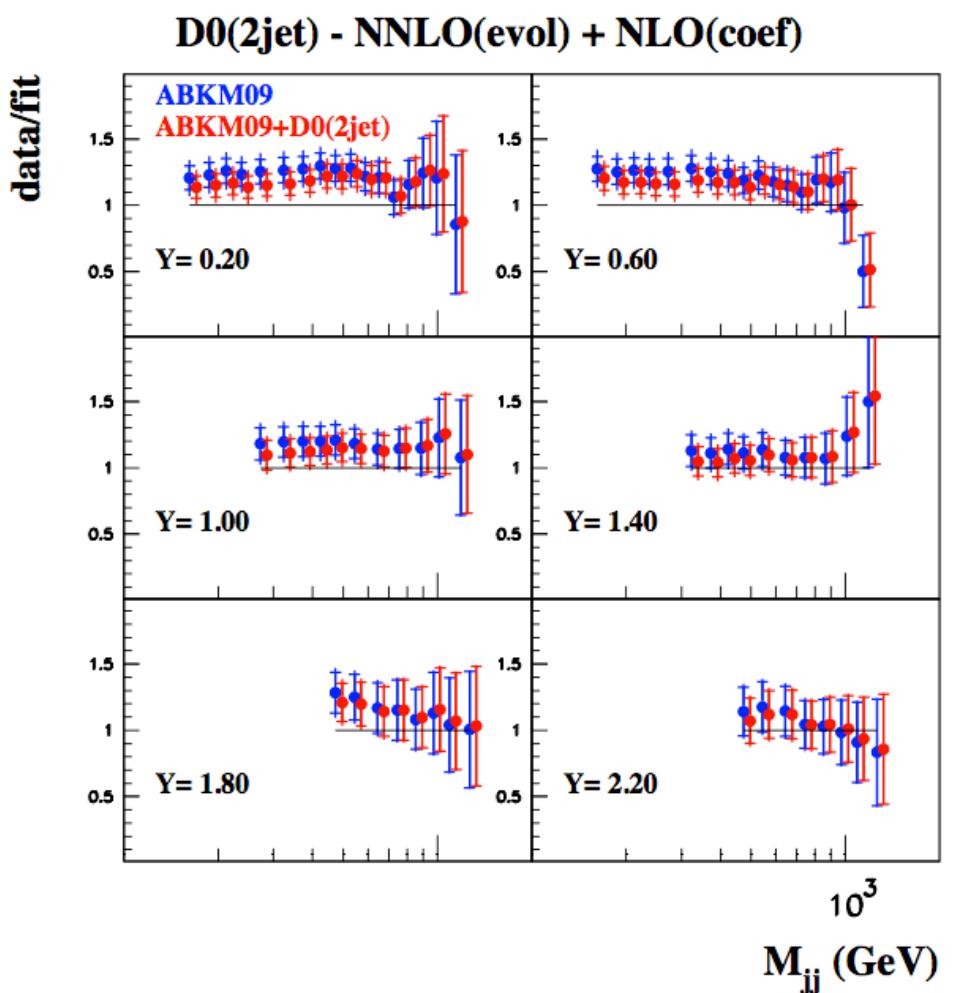
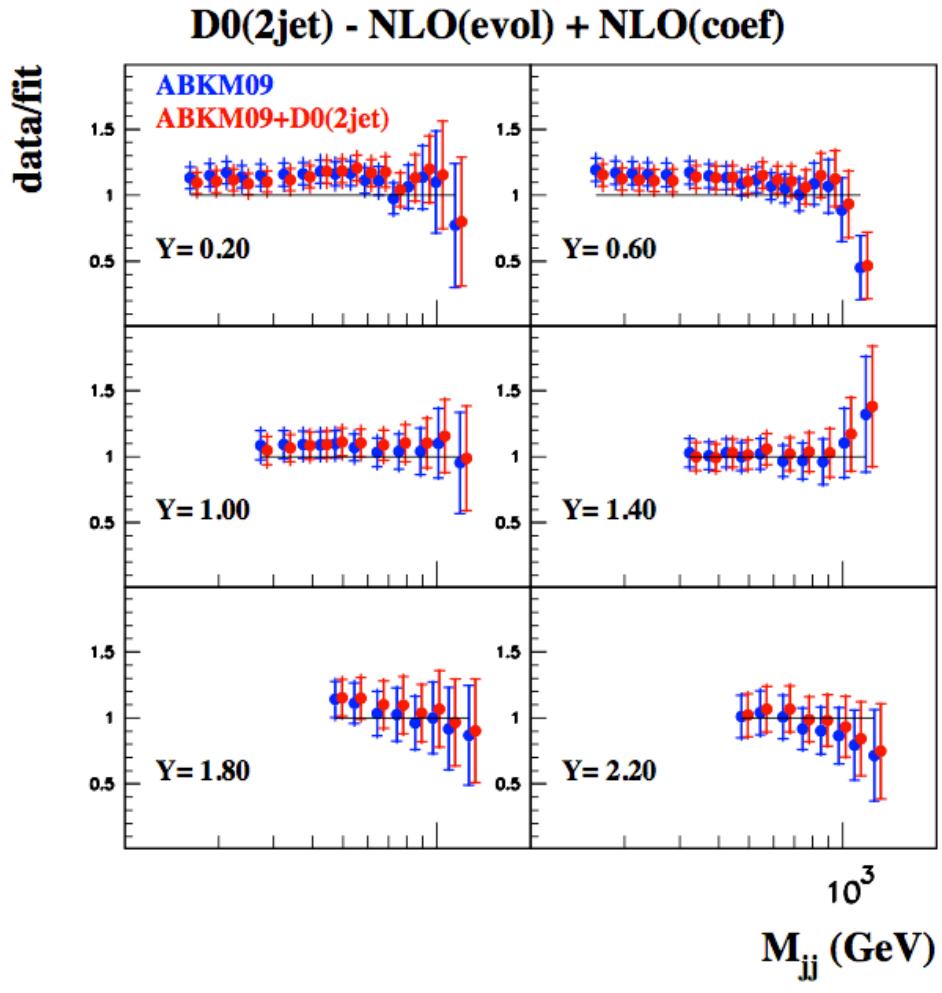


$m_H = 165$  GeV Tevatron Higgs production cross sections from ABKM

$\sigma(H)$	ABKM PDFs	1-jet inclusive data	di-jet data
NLO	0.206(17) pb	0.235(10) pb	0.212(9) pb
NNLO	<b>0.253(22) pb</b>	0.297(12) pb	0.278(13) pb

c.f. NNLO with MSTW08: 0.341 pb, HERAPDF: 0.269 pb

# ABKM09 Fitted with D0 Run II Dijet $m_{jj}$ Data



# MSTW08 Fits to CDF and D0 run II Jet Data

**CDF Run II inclusive jet data,  $\chi^2 = 56$  for 76 pts.**

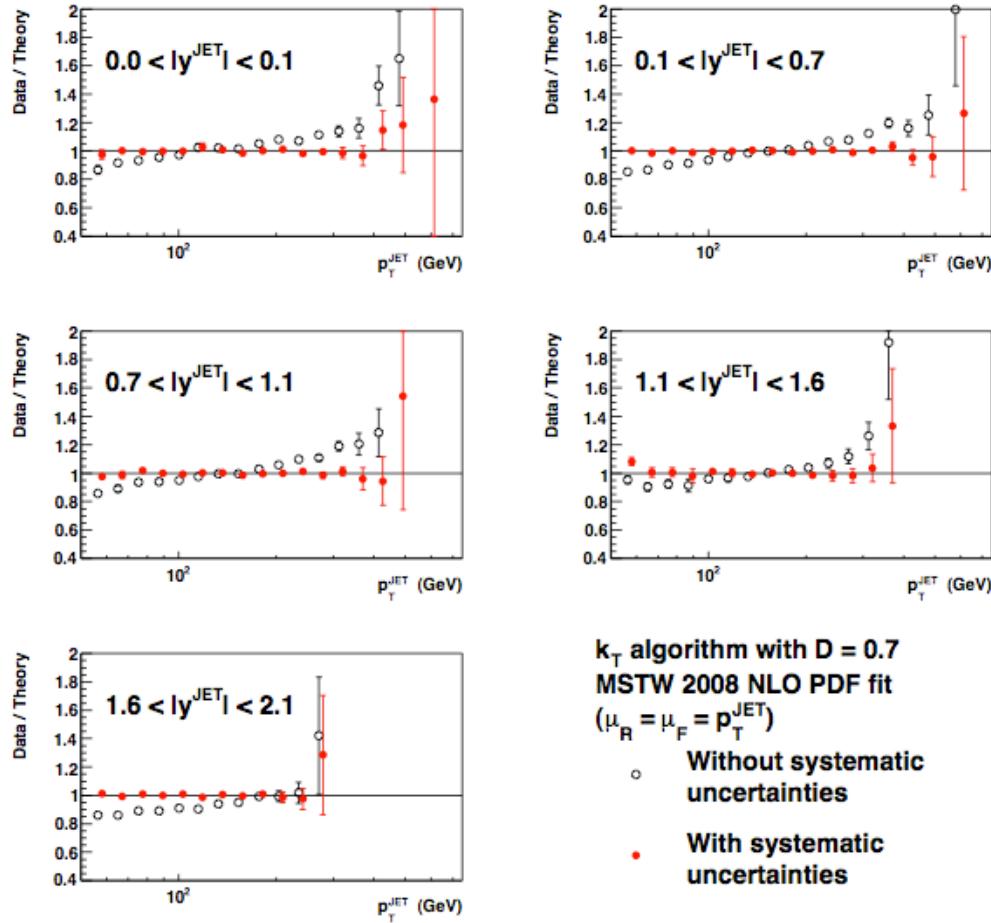
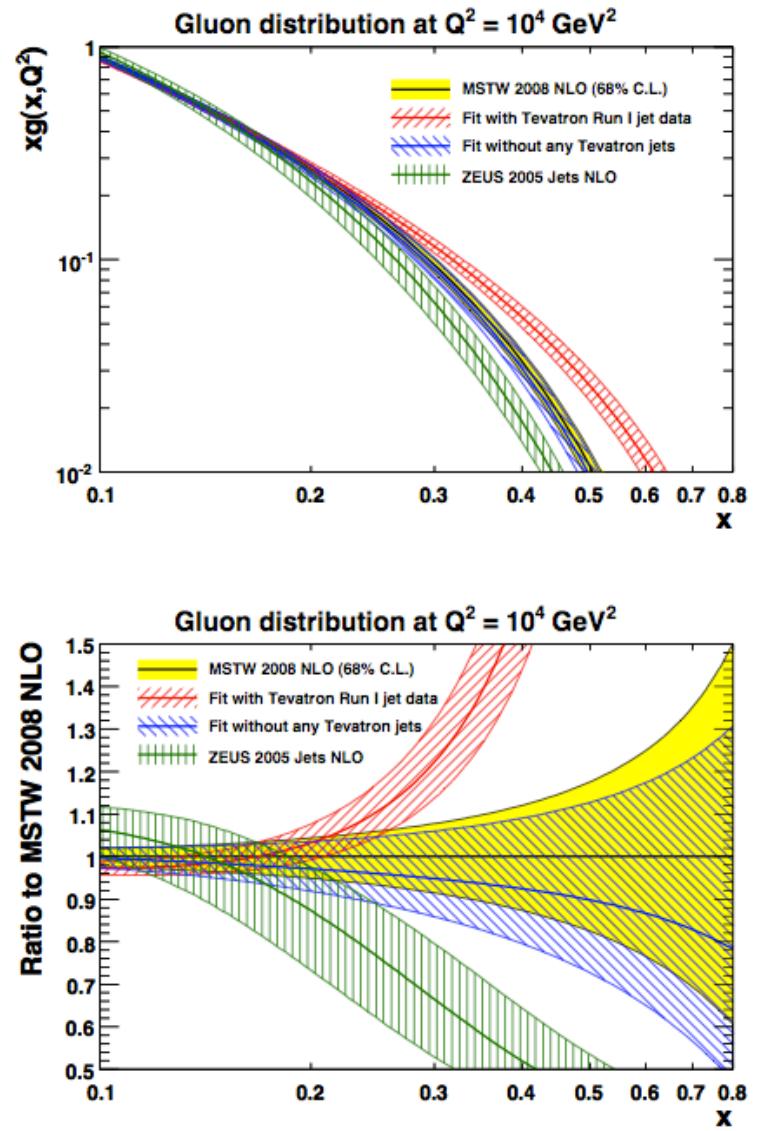


Figure 48: Data/theory ratio for the CDF Run II data obtained using the  $k_T$  jet clustering algorithm [54] and the MSTW 2008 NLO fit.



## DØ Run II inclusive jet data (cone, R = 0.7)

MSTW 2008 NLO PDF fit ( $\mu_R = \mu_F = p_T^{\text{JET}}$ ),  $\chi^2 = 114$  for 110 pts.

Jet production predicted with FastNLO; MSTW estimate NNLO corrections to be small. Errors incurred in omitting this data are expected to be much bigger than those incurred by using only NLO predictions in making NNLO PDFs.

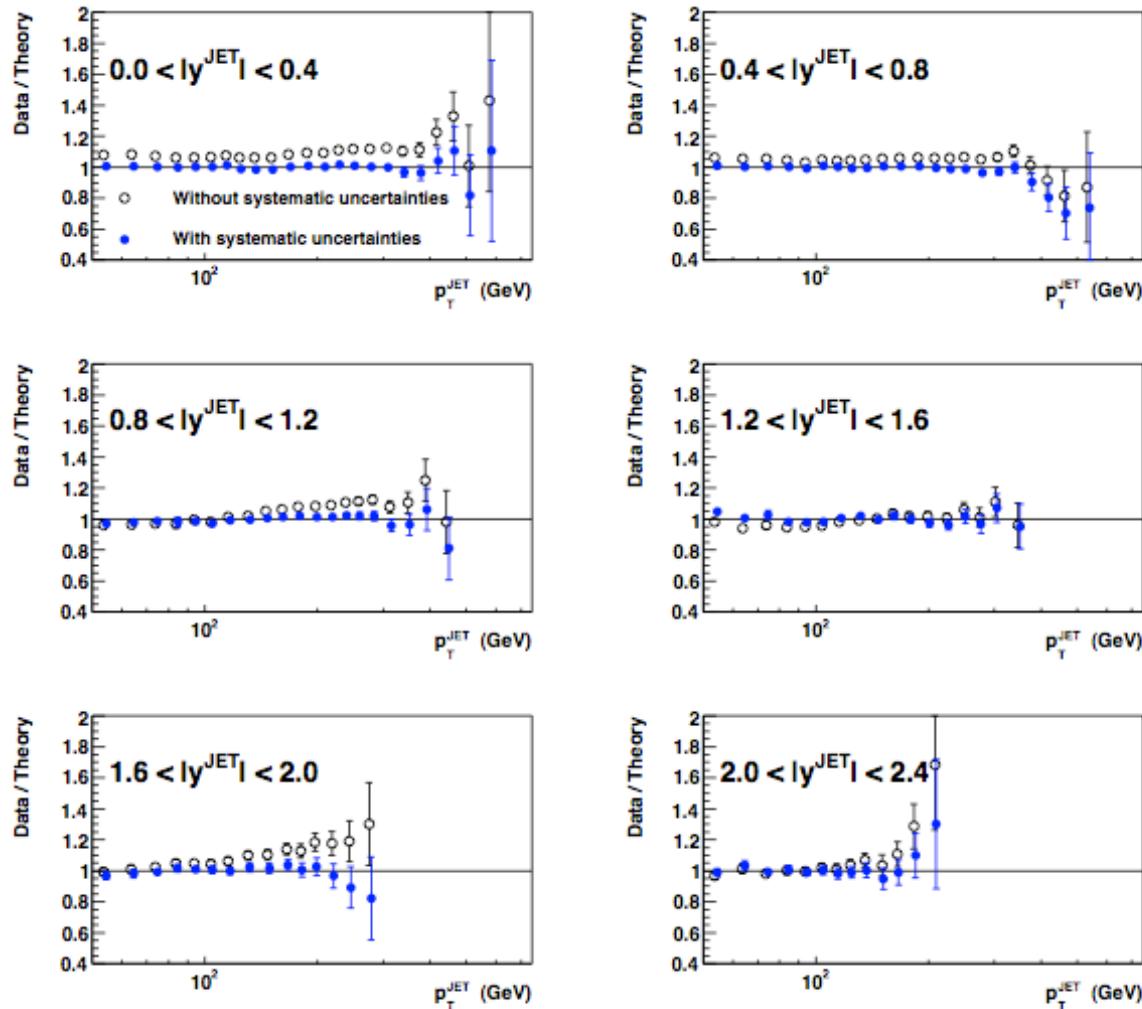
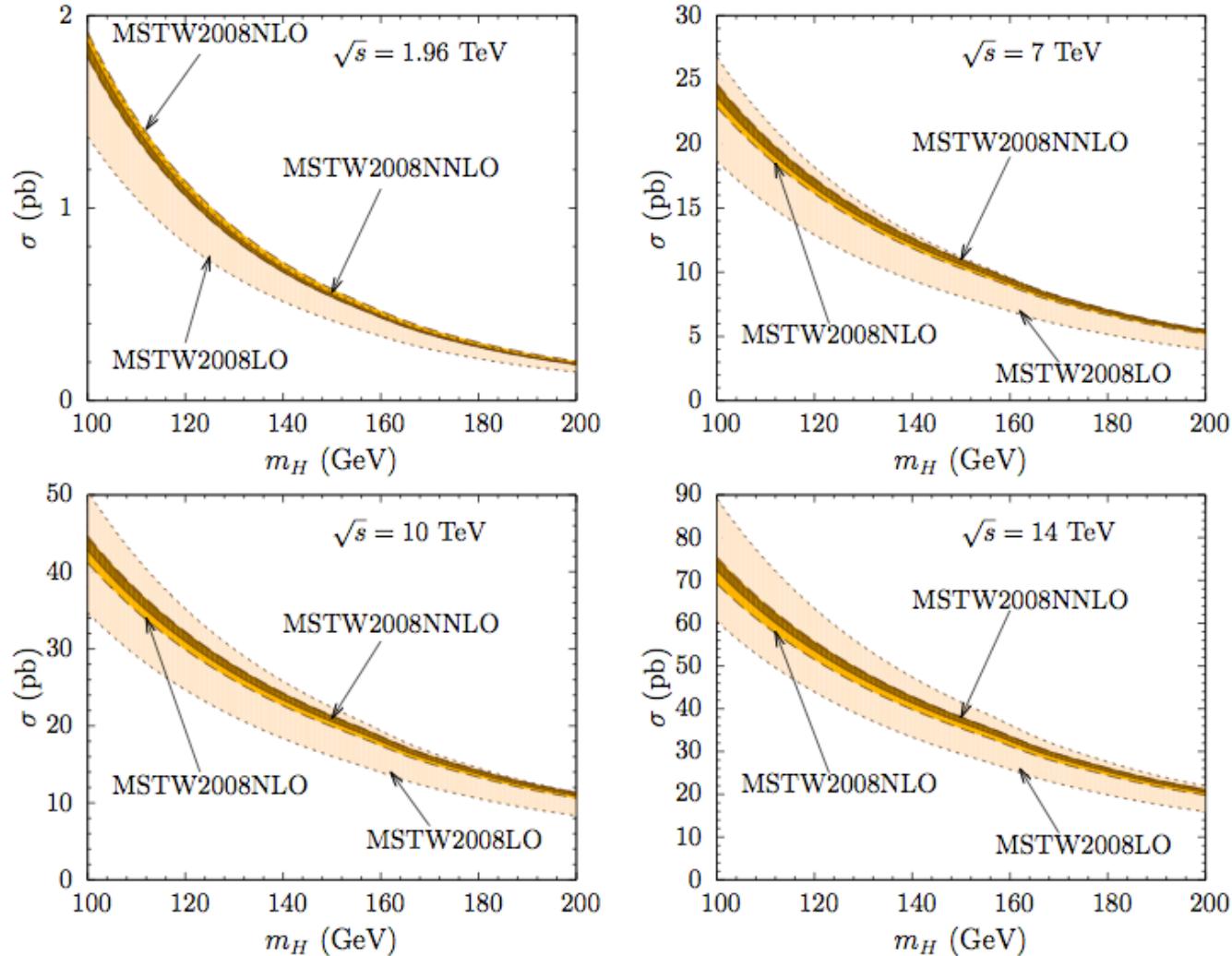


Figure 49: Data/theory ratio for the DØ Run II data [56] and the MSTW 2008 NLO fit.



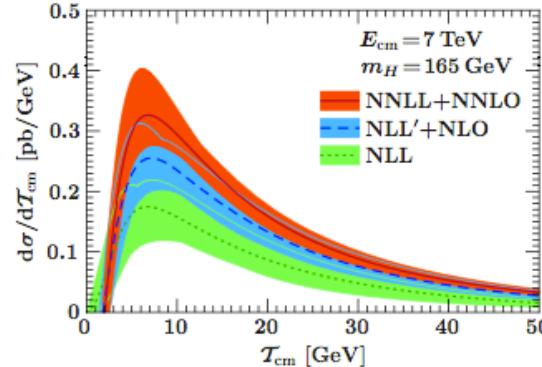
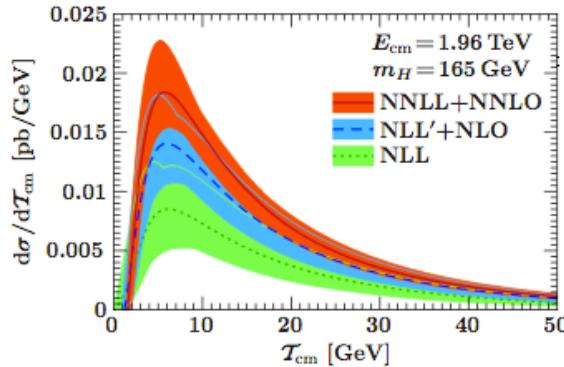
Rapid convergence  
of perturbative  
series with  
resummation.

Figure 1: Cross sections at the Tevatron for  $\sqrt{s} = 1.96$  TeV and the LHC for  $\sqrt{s} = 7, 10, 14$  TeV. Bands indicate scale uncertainties. Light, medium and dark bands represent LO (NLL), NLO (NNLL) and NNLO ( $N^3LL$ ) in RG-improved perturbation theory, respectively.

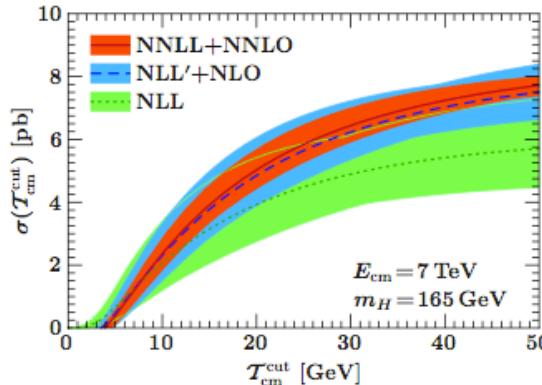
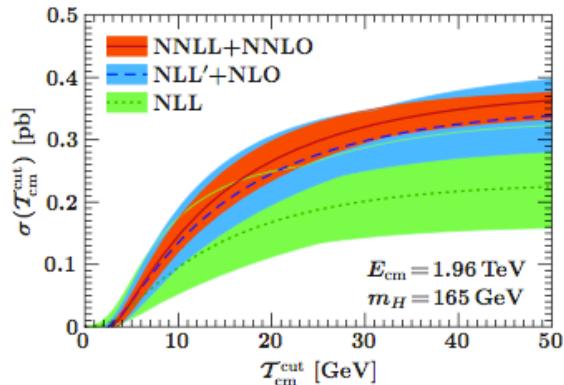
# A Stand-In for a Jet Veto that can be Calculated at NNLO+NNLL

$$\tau = \frac{\mathcal{T}_{\text{cm}}}{m_H}, \quad \mathcal{T}_{\text{cm}} = \sum_k |\vec{p}_{kT}| e^{-|\eta_k|} = \sum_k (E_k - |p_k^z|).$$

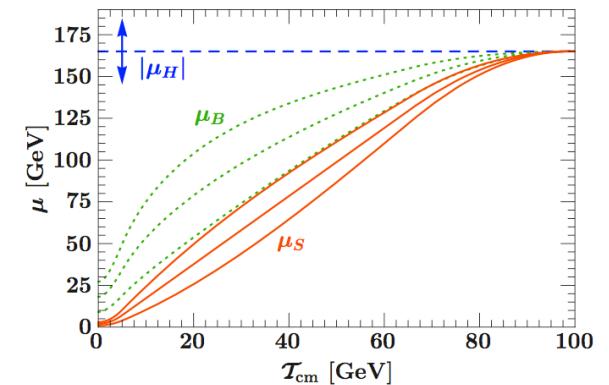
Low Beam thrust means no jets.  
(converse?)



**Figure 9.** The beam thrust spectrum for Higgs production for  $m_H = 165$  GeV at the Tevatron (left) and the LHC for  $E_{\text{cm}} = 7$  TeV (right). The bands show the perturbative scale uncertainties as explained in section 2.6.

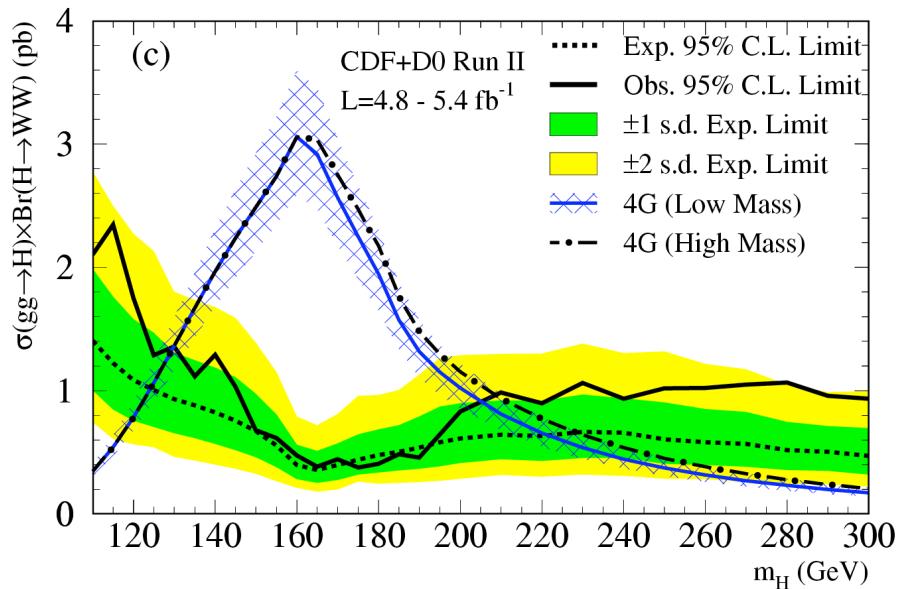


**Figure 10.** Higgs production cross section as a function of  $\mathcal{T}_{\text{cm}}^{\text{cut}}$  for  $m_H = 165$  GeV at the Tevatron (left) and the LHC with  $E_{\text{cm}} = 7$  TeV (right). The bands show the perturbative scale uncertainties as explained in section 2.6.



Berger, Marcantonini,  
Stewart, Tackmann, and  
Waalwijn, arXiv:1012.4480

# Setting a Limit on $\sigma(gg \rightarrow H) \times Br(H \rightarrow W^+W^-)$



[T. Aaltonen et al., Phys. Rev. D 82, 011102\(R\) \(2010\)](#)

- Model-”Independent”!
- No total uncertainty on the cross section or b.r. from PDF or scale needed when setting this limit.
- But: We need acceptance for our selection! 0j, 1j, 2+ jets have relative predictions that depend on scale (and less on PDF)

Our procedure: take the jet-category uncertainties and subtract off the common total amount from each one, so that we preserve the uncertainties in the jet categories while removing the total uncertainty.

# Cross Section and Branching Ratio Alignment

Table 1: The NNLO production cross section for  $WH$ ,  $ZH$ , and the associated uncertainties

## WH and ZH Production rates at the Tevatron

Baglio and Djouadi  
JHEP 1010, 064 (2010)  
arXiv:1003.4266v2

Updated from v1:  
Proper CKM elements in WH  
More precise integrations  
Reductions of cross sections  
of 4-6%

$m_H$ (GeV/c <sup>2</sup> )	$\sigma_{WH}$ (fb)	$\sigma_{ZH}$ (fb)	scale (%)	PDF+ $\alpha_s^{\text{exp}}$ (%)	$\alpha_s^{\text{th}}$ (%)
100	291.90	169.8	+0.7 -0.8	+5.0 -5.0	+0.5 -0.4
105	248.40	145.9	+0.7 -0.8	+5.4 -5.4	+0.6 -0.4
110	212.00	125.7	+0.7 -0.9	+5.8 -5.8	+0.6 -0.5
115	174.50	103.9	+0.7 -0.9	+6.1 -6.1	+0.7 -0.5
120	150.10	90.2	+0.7 -0.9	+6.4 -6.3	+0.8 -0.6
125	129.50	78.5	+0.7 -1.0	+6.6 -6.7	+0.8 -0.6
130	112.00	68.5	+0.7 -1.0	+6.4 -6.7	+1.0 -0.7
135	97.20	60.0	+0.7 -1.0	+6.9 -6.8	+1.0 -0.7
140	84.60	52.7	+0.7 -1.1	+6.9 -6.7	+1.1 -0.8
145	73.70	46.3	+0.7 -1.1	+7.3 -7.1	+1.2 -0.9
150	64.40	40.8	+0.8 -1.1	+6.8 -6.7	+1.2 -0.9
155	56.20	35.9	+0.7 -1.1	+7.5 -7.3	+1.2 -1.1
160	48.50	31.4	+0.8 -1.2	+7.4 -6.8	+1.4 -1.0
165	43.60	28.4	+0.7 -1.1	+7.8 -7.6	+1.4 -1.1
170	38.50	25.3	+0.8 -1.0	+7.8 -7.0	+1.6 -1.3
175	34.00	22.5	+0.9 -1.2	+7.9 -7.6	+1.5 -1.2
180	30.10	20.0	+0.7 -1.3	+7.3 -7.3	+1.7 -1.3
185	26.90	17.9	+0.7 -1.1	+7.8 -7.8	+1.9 -1.5
190	24.00	16.1	+0.8 -1.2	+7.5 -7.5	+1.7 -1.2
195	21.40	14.4	+1.4 -1.4	+8.4 -7.9	+1.9 -1.4
200	19.10	13.0	+1.0 -1.0	+7.9 -7.3	+2.1 -1.6
210	15.20	10.5	+0.7 -0.7	+7.2 -7.2	+1.3 -1.3
220	12.30	8.5	+0.3 -0.3	+7.4 -7.4	+1.1 -1.3
230	9.90	7.0	+0.4 -0.4	+7.6 -7.6	+1.2 -1.4
240	8.03	5.7	+0.2 -0.2	+7.7 -7.7	+1.4 -1.5
250	6.53	4.7	+0.3 -0.3	+7.8 -8.0	+1.4 -1.5
260	5.33	3.9	+0.2 -0.2	+8.1 -8.1	+1.5 -1.7
270	4.37	3.2	+0.2 -0.2	+8.0 -8.2	+1.6 -1.6
280	3.59	2.7	+0.3 -0.3	+8.1 -8.4	+1.7 -1.7
290	2.96	2.2	+0.3 -0.3	+8.4 -8.8	+1.7 -1.7
300	2.45	1.9	+0.0 -0.0	+8.6 -8.6	+1.6 -2.0

# Cross Section and Branching Ratio Alignment

## Weak Vector Boson Fusion (VBF)

NNLO Calc from  
P. Bolzoni, F. Maltoni,  
S.-O. Moch, and M. Zaro

<http://vbf-nnlo.phys.ucl.ac.be/vbf.html>

Very close to the NLO calculation.

We checked the NLO EW corrections from  
HAWK: Denner, Dittmaier, Mück  
and found that they are of order -3%

$m_H$ (GeV/c <sup>2</sup> )	$\sigma_{\text{VBF}}$ (NNLO) (fb)	scale (%)	PDF+ $\alpha_s$ (%)	$\sigma_{\text{VBF}}$ (HAWK) (fb)
100	100.1	+0.7 -0.3	$\pm 4.8$	98.6
105	92.3	+0.7 -0.3	$\pm 4.9$	90.8
110	85.1	+0.8 -0.3	$\pm 4.9$	83.8
115	78.6	+0.8 -0.3	$\pm 4.9$	77.3
120	72.7	+0.6 -0.4	$\pm 4.9$	71.4
125	67.1	+0.7 -0.3	$\pm 5.0$	66.0
130	62.1	+0.7 -0.4	$\pm 5.0$	61.0
135	57.5	+0.7 -0.4	$\pm 5.0$	56.5
140	53.2	+0.7 -0.4	$\pm 5.0$	52.3
145	49.4	+0.7 -0.4	$\pm 5.1$	48.5
150	45.8	+0.7 -0.5	$\pm 5.1$	44.9
155	42.4	+0.7 -0.5	$\pm 5.1$	41.5
160	39.4	+0.7 -0.5	$\pm 5.1$	38.5
165	36.6	+0.7 -0.5	$\pm 5.1$	36.2
170	34.0	+0.7 -0.6	$\pm 5.2$	33.6
175	31.6	+0.7 -0.6	$\pm 5.2$	31.2
180	29.4	+0.7 -0.6	$\pm 5.2$	28.9
185	27.3	+0.7 -0.6	$\pm 5.2$	27.0
190	25.4	+0.7 -0.6	$\pm 5.2$	25.1
195	23.7	+0.8 -0.7	$\pm 5.3$	23.4
200	22.0	+0.8 -0.7	$\pm 5.3$	21.7
210	19.1	+0.7 -0.7	$\pm 5.3$	18.8
220	16.6	+0.7 -0.8	$\pm 5.3$	16.3
230	14.5	+0.8 -0.8	$\pm 5.4$	14.2
240	12.6	+0.8 -0.8	$\pm 5.4$	12.3
250	11.0	+0.8 -0.9	$\pm 5.4$	10.7
260	9.6	+0.8 -1.0	$\pm 5.5$	9.4
270	8.4	+0.8 -1.0	$\pm 5.5$	8.2
280	7.4	+0.8 -1.1	$\pm 5.5$	7.1
290	6.4	+0.8 -1.1	$\pm 5.6$	6.2
300	5.6	+0.8 -1.1	$\pm 5.6$	5.4

# Tevatron Correlated Systematic Errors I

Total Systematic error count: 129 (not counting bin-by-bin errors)

Note: correlation in errors on backgrounds between experiments helps sensitivity! One experiment is another experiment's control sample.

Luminosity: 3.8% Correlated CDF and D0  $\sigma_{\text{inel}}(\text{ppbar})$   
4.4% detector-specific

Diboson Cross Sections:

For WZ and ZZ, require  
 $75 < m_{\parallel} < 105 \text{ GeV}$

$$\sigma_{W^+W^-} = 11.34^{+0.56}_{-0.49} \text{ (scale)} \quad {}^{+0.35}_{-0.28} \text{ (PDF) pb}$$

$$\sigma_{W^\pm Z} = 3.22^{+0.20}_{-0.17} \text{ (scale)} \quad {}^{+0.11}_{-0.08} \text{ (PDF) pb}$$

$$\sigma_{ZZ} = 1.20^{+0.05}_{-0.04} \text{ (scale)} \quad {}^{+0.04}_{-0.03} \text{ (PDF) pb}$$

WW, WZ, and ZZ total cross section uncertainties considered 100% correlated with each other (Joey Huston computed PDF correlations and found they are nearly 100%). We measure the diboson background *in situ* with similar precision to the prediction.

ttbar Cross Section: Moch and Uwer, evaluated at (also measured *in situ*)  
 $m_t = 173 \pm 1.2 \text{ GeV}$  is (using MSTW2008 PDFs),

$$\sigma_{t\bar{t}} = 7.04^{+0.24}_{-0.36} \text{ (scale)} \pm 0.14 \text{ (PDF)} \pm 0.30 \text{ (mass)}$$

# Tevatron Correlated Systematic Errors II

Signal Cross Section uncertainties (using MSTW PDFs)

WH, ZH:  $\pm 5\%$

$gg \rightarrow H$ :  $\pm \sim 20\%$  (weighted scale over jet samples)  
 $\pm 12.5\%$  (weighted PDF). Pt spectra reweighted to NNLO+NNLL predictions

Errors taken from Anastasiou, Grazzini, Dissertori, Stockli, and Webber, for H+0, 1 jets

VBF:  $\pm 10\%$

Theory errors applied to SM interpretations, but taken off for cross-section times branching ratio limits.

*CDF-D0 Uncorrelated errors:*

- K-factors (data driven)
- trigger efficiency
- b-tag efficiency and mistags
- jet energy scale
- lepton ID, fakes and conversions
- MET modeling



Correlated *within* CDF and D0 where appropriate

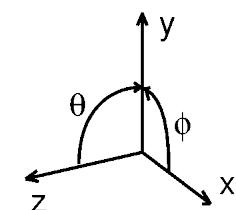
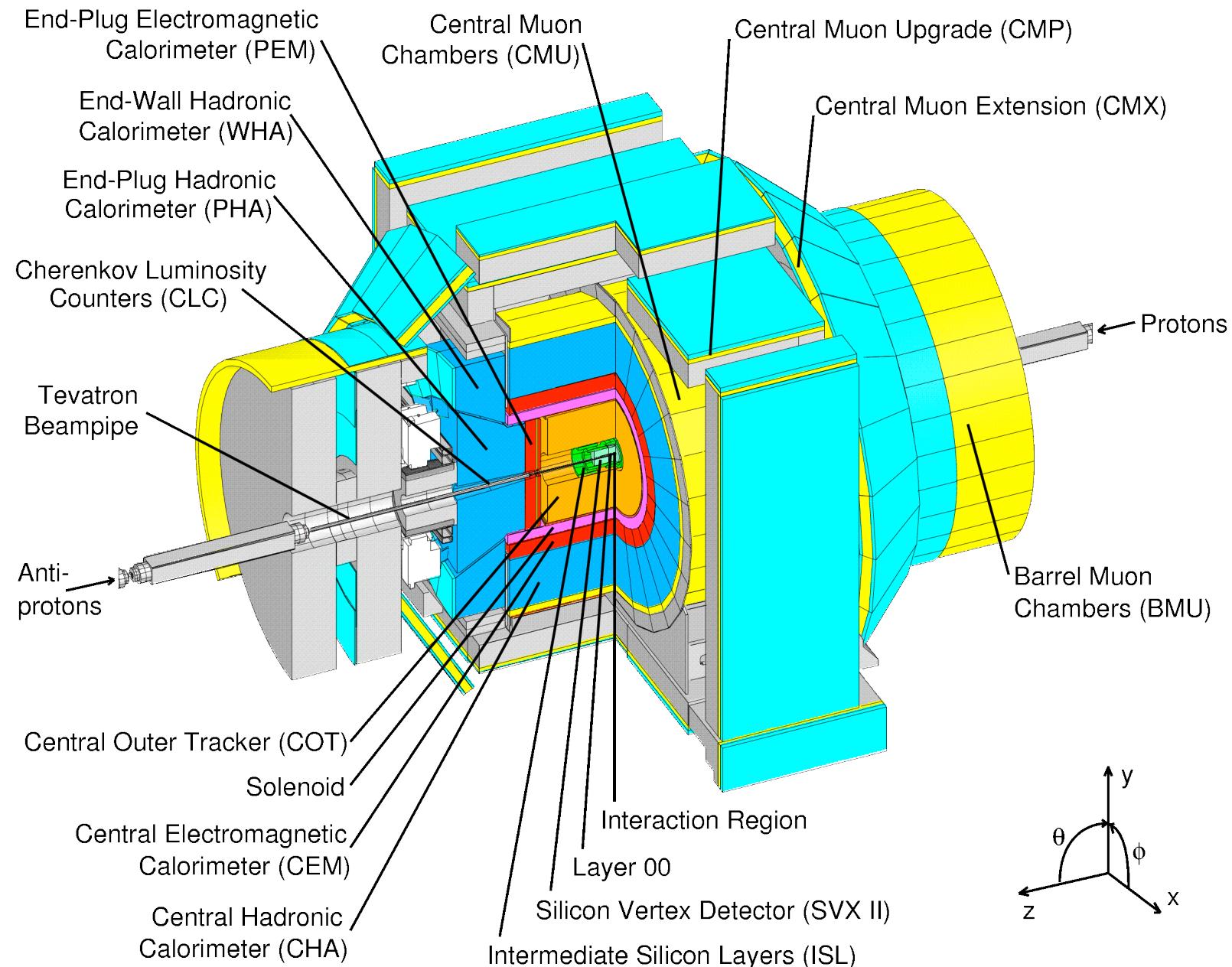
# Including Higher-Order Calculations

We like theoretical work which addresses our selections and use the resulting systematic uncertainties.

But we realize that the mapping between partons and measurements is not perfect.

- showering, hadronization (available to everyone)
- Detector acceptance, efficiency, resolution (not available to everyone)
- Need fully simulated Monte Carlo – happy to adjust the underlying physics of the generator though.

# The CDF Detector



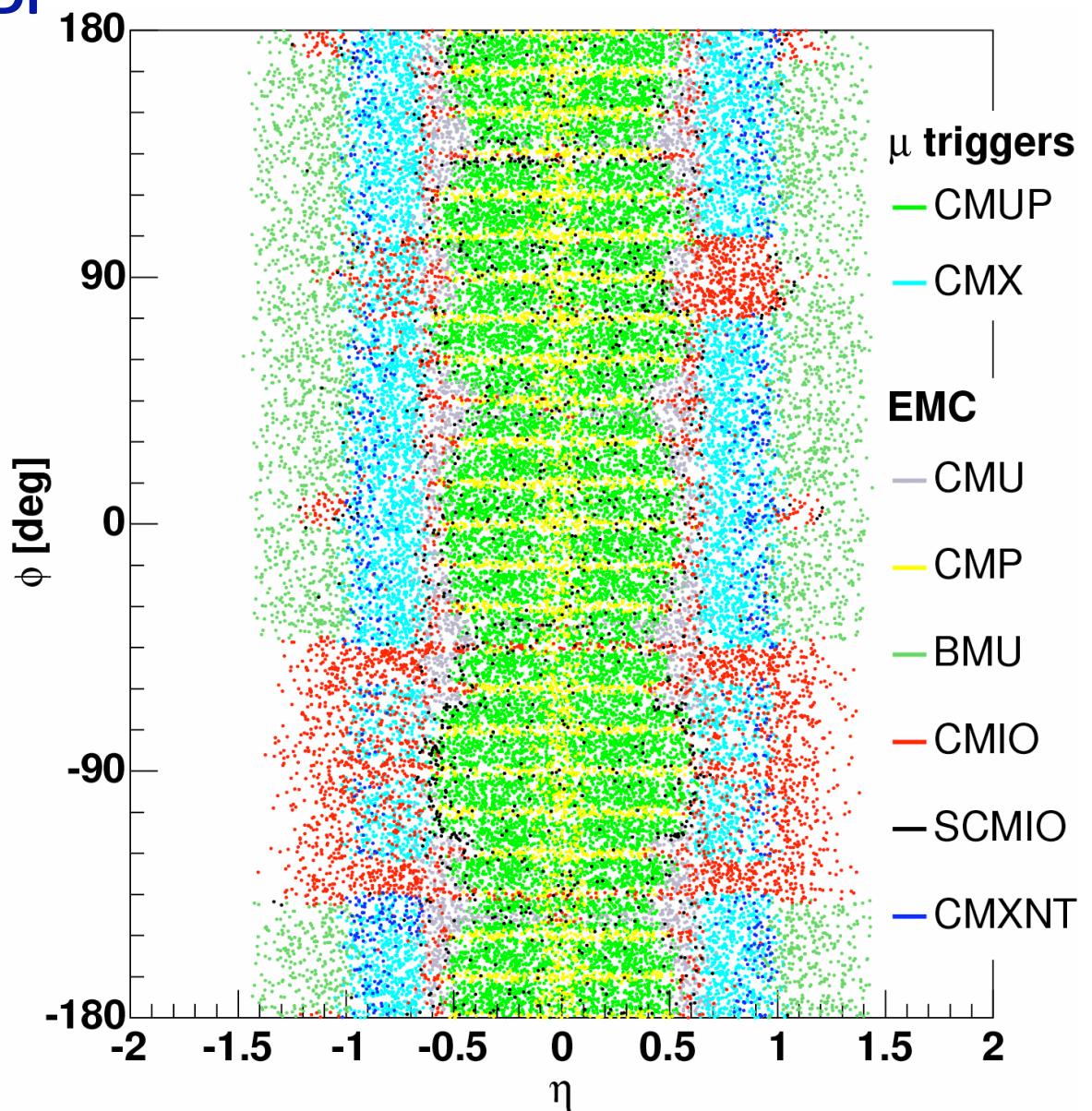
## Muon Detection in CDF

requires looking at many different detectors and categories of muons.

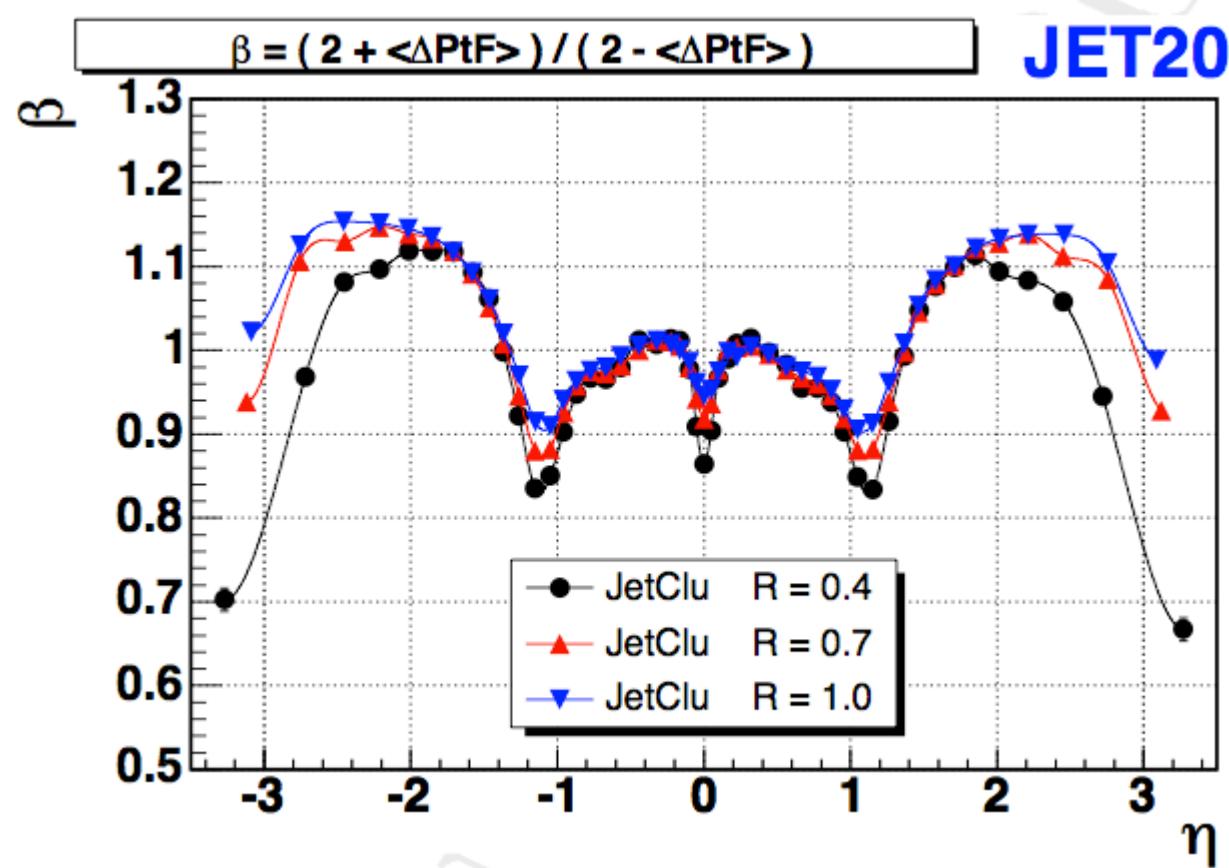
Detailed detector simulation + data-calibrated identification rate scale factors and resolutions required.

Calibrated with  $Z \rightarrow \mu\mu$

Latest iteration have even more categories  
“crack tracks”



# Uncorrected Jet Response is Highly angle dependent



This is corrected out, but then the jet resolution is highly angle dependent.

# Steps Required for Combination

Before Exchanging data, quality control:

- Check stacked histograms and systematic tables with analysis documentation total counts:  
data, signal, background
  - look for bins with  $b=0$  and have data events (bad!)
- Repeat individual channel limits -- compare against approved results.

CDF and D $\emptyset$  teams each do three combinations, using Bayesian and CL<sub>s</sub> techniques.

CDF

D $\emptyset$

Tevatron

Consistency at the better than 10% level required for all combinations at all test masses. Quote Bayesian limits (historical)

# Mini-Review: Bayesian Limits

$$L(r, \theta) = \prod_{\text{channels}} \prod_{\text{bins}} P_{\text{Poiss}}(\text{data} | r, \theta)$$

Where  $r$  is an overall signal scale factor, and  $\theta$  represents all nuisance parameters.

$$P_{\text{Poiss}}(\text{data} | r, \theta) = \frac{(rs_i(\theta) + b_i(\theta))^{n_i} e^{-(rs_i(\theta) + b_i(\theta))}}{n_i!}$$

where  $n_i$  is observed in each bin  $i$ ,  $s_i$  is the predicted signal for a fiducial model (SM), and  $b_i$  is the predicted background. Dependence of  $s_i$  and  $b_i$  on  $\theta$  includes rate, shape, and bin-by-bin independent uncertainties.

Tests models that scale (unphysically) WH, ZH, VBF, ggH all together, holding decay b.r.'s fixed

# Mini-Review: Bayesian Limits

Including uncertainties on nuisance parameters  $\theta$

$$L'(data | r) = \int L(data | r, \theta) \pi(\theta) d\theta$$

- $\pi(\theta)$  encodes our prior belief in the values of the uncertain parameters.
- Nuisance parameters have correlated impacts on multiple predictions.
- Integral is a convolution for uncertainties that add: “quadrature” for uncorrelated uncertainties, “linear” for correlated uncertainties.
- Some uncertainties are multiplicative, and their impacts are propagated as such to the predictions.
- Rate, shape, bin-by-bin uncertainties handled together

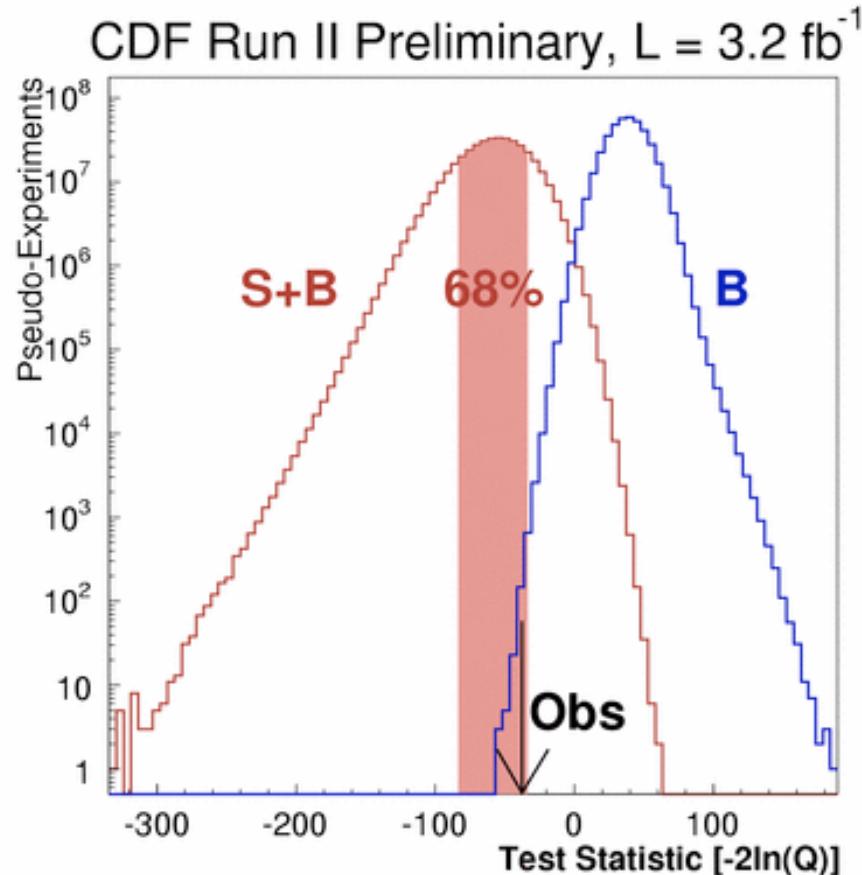
Useful for a variety of results:

Limits:  $0.95 = \int_0^{r_{\lim}} L'(data | r) \pi(r) dr$

Typically  $\pi(r)$  is constant  
Other options possible.  
Sensitivity to priors a concern.

Measure  $r$ :  $0.68 = \int_{r_{low}}^{r_{high}} L'(data | r) \pi(r) dr$     $r = r_{\max} \frac{+(r_{high} - r_{\max})}{(r_{\max} - r_{low})}$

# Discovery with p-Values



Example: CDF single top.

$$-2\ln Q \equiv LLR \equiv -2\ln \left( \frac{L(\text{data} | s+b, \theta)}{L(\text{data} | b, \hat{\theta})} \right)$$

100 M s+b and b-only pseudoexperiments, each with fluctuated nuisance parameters, and fit twice.

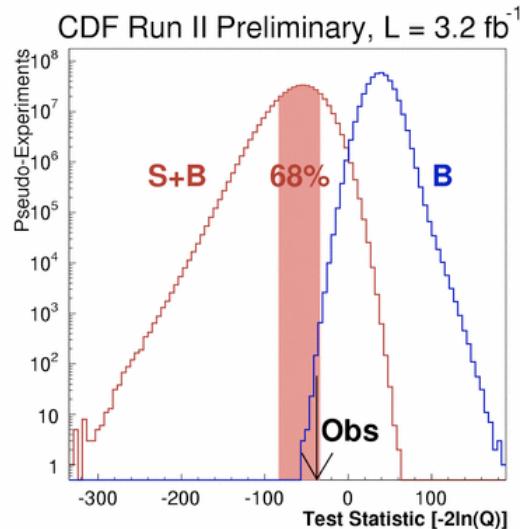
5 $\sigma$ : p-value of  $2.77 \times 10^{-7}$  or less.

3 $\sigma$ : p-value of  $1.35 \times 10^{-3}$  or less

2 $\sigma$ : p-value of 2.28% or less

Buzzword: "Prior Predictive ensemble"

# Fitting and Fluctuating



$$-2\ln Q \equiv LLR \equiv -2\ln \left( \frac{L(\text{data} | s + b, \hat{\theta})}{L(\text{data} | b, \hat{\theta})} \right)$$

- Monte Carlo pseudoexperiments are used to get p-values.
- Test statistic  $-2\ln Q$  is not uncertain for the data.
- Distribution from which  $-2\ln Q$  is drawn is uncertain!

- Nuisance parameter fits in numerator and denominator of  $-2\ln Q$  **do not incorporate systematics into the result.**  
Example -- 1-bin search; all test statistics are equivalent to the event count, fit or no fit.
- Instead, we fluctuate the probabilities of getting each outcome since those are what we do not know. Each pseudoexperiment gets random values of nuisance parameters.
- Can also try values of nuisance parameters that maximize the p-value, but that's very conservative (called the supremum p-value, still needs choices of parameter ranges).
- Why fit at all? It's an optimization. Fitting reduces sensitivity to the uncertain true values and the fluctuated values. For stability and speed, you can choose to fit a subset of nuisance parameters (the ones that are constrained by the data). Or do constrained or unconstrained fits, it's your choice.
- If not using pseudoexperiments but using Wilk's theorem, then the fits are important for correctness, not just optimality.

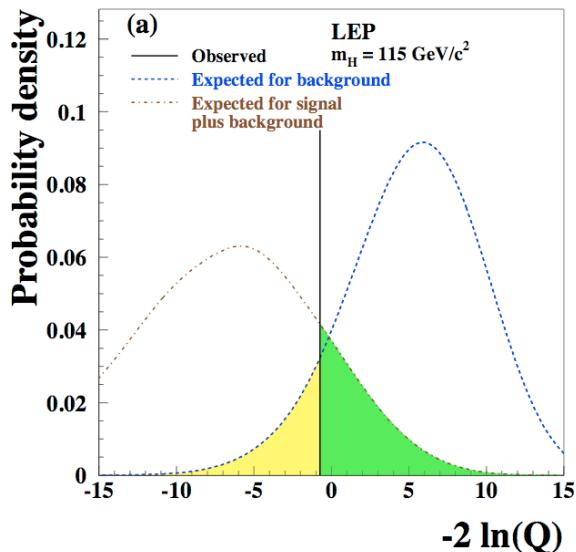
## Mini-Review: $CL_s$ Limits

- Based on p-values using the log likelihood ratio as the test statistic. Neyman-Pearson lemma says LLR is the uniformly most powerful test statistic, although the Neyman-Pearson one fits for the parameter of interest, not just the nuisance parameters, making the null hypothesis a subset of the test hypothesis

$$-2\ln Q \equiv LLR \equiv -2\ln \left( \frac{L(\text{data} | s + b, \hat{\theta})}{L(\text{data} | b, \hat{\theta})} \right)$$

Pearson's LLR also fits for  $s$  (actually  $r \times s$ ) in the numerator, while  $r = 0$  in the denominator

# Mini-Review: $CL_s$ Limits



p-values:

$$\text{Yellow area} = 1 - CL_b = 1 - P(-2 \ln Q > -2 \ln Q_{\text{obs}} \mid \text{b only})$$

$$\text{Green area} = CL_{s+b} = P(-2 \ln Q > -2 \ln Q_{\text{obs}} \mid s+b)$$

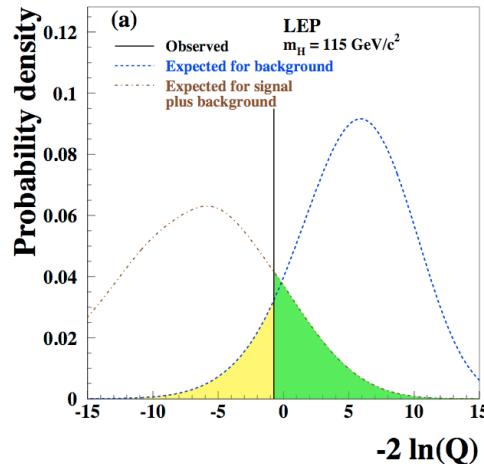
$$CL_s = CL_{s+b} / CL_b \geq CL_{s+b}$$

Exclude if  $CL_s < 0.05$

Vary  $r$  until  $CL_s = 0.05$  to get  $r_{\text{lim}}$

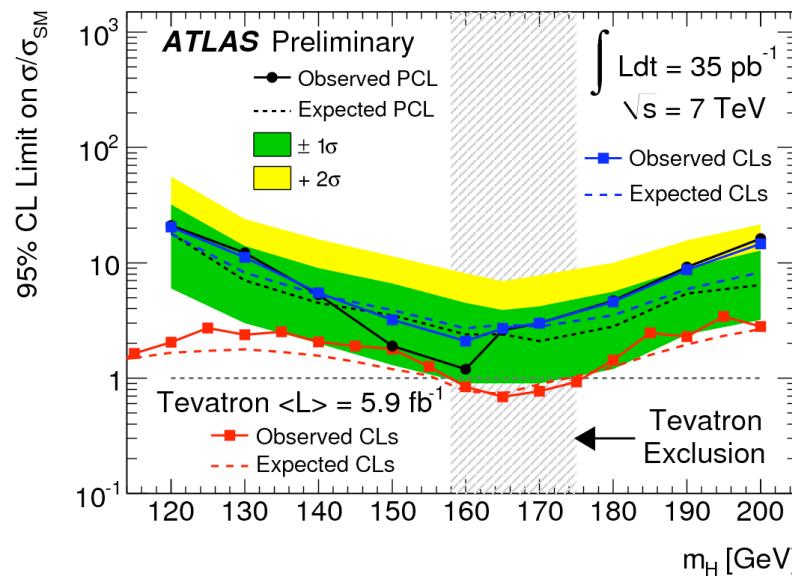
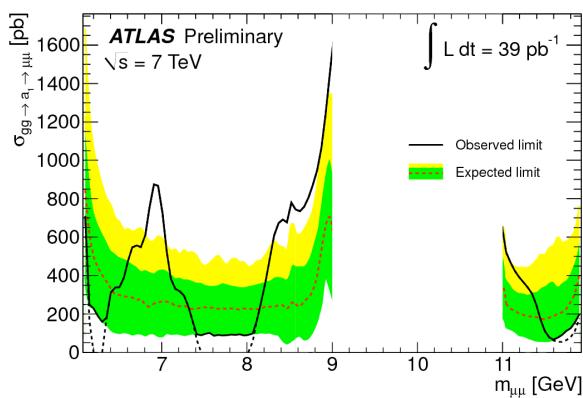
- Advantages:
  - Exclusion and Discovery p-values are consistent.  
Example -- a  $2\sigma$  upward fluctuation of the data with respect to the background prediction appears both in the limit and the p-value as such
  - Does not exclude where there is no sensitivity  
(big enough search region with small enough resolution and you get a 5% dusting of random exclusions with  $CL_{s+b}$ )

# Power Constrained Limits (PCL)



Just use  $CL_{S+b} < 0.05$  to determine exclusion.

But if the resulting limit is more than  $1\sigma$  more stringent than the median expectation, quote the  $1\sigma$  limit instead

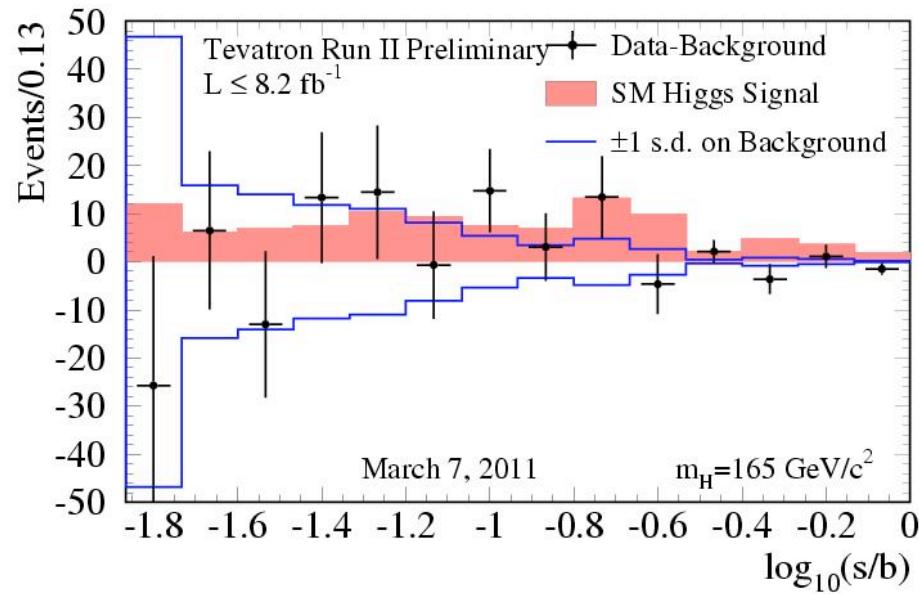
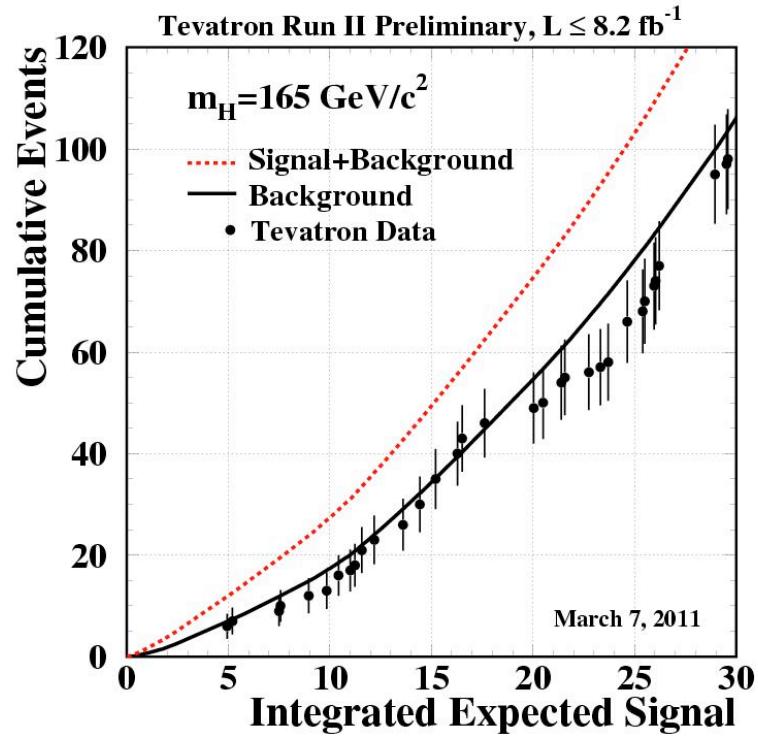
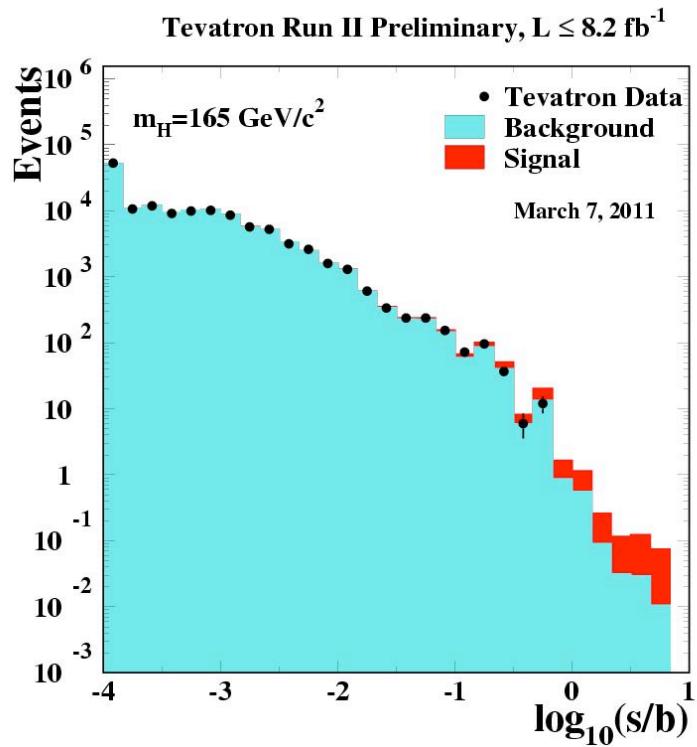


Advantages:

- More powerful than  $CL_s$  or Bayesian limits while still covering
- Does not exclude where there is no sensitivity

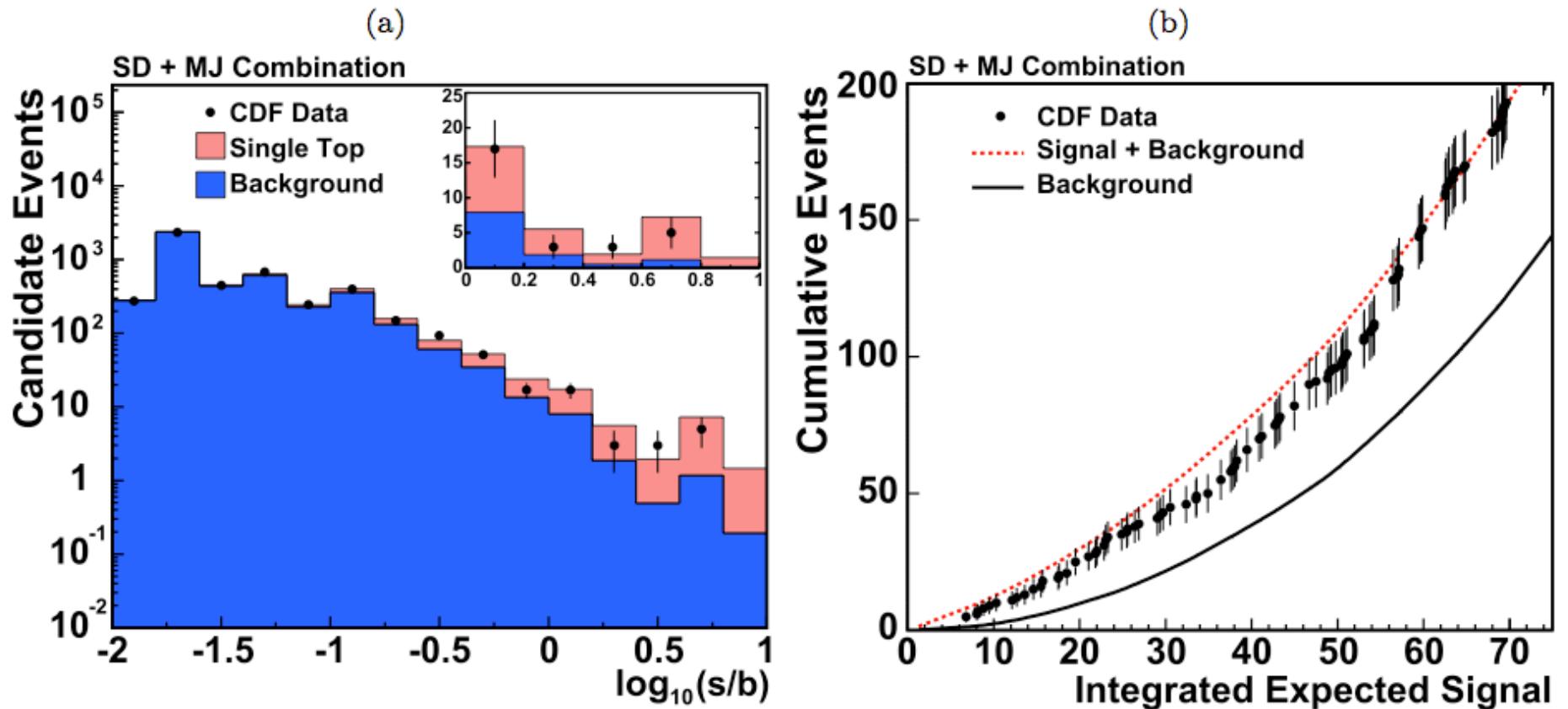
Disadvantage:

- $1\sigma$  constraint is arbitrary – balance desire for a more powerful method with acceptability of limits. A  $2\sigma$  constraint defeats the purpose entirely for example.



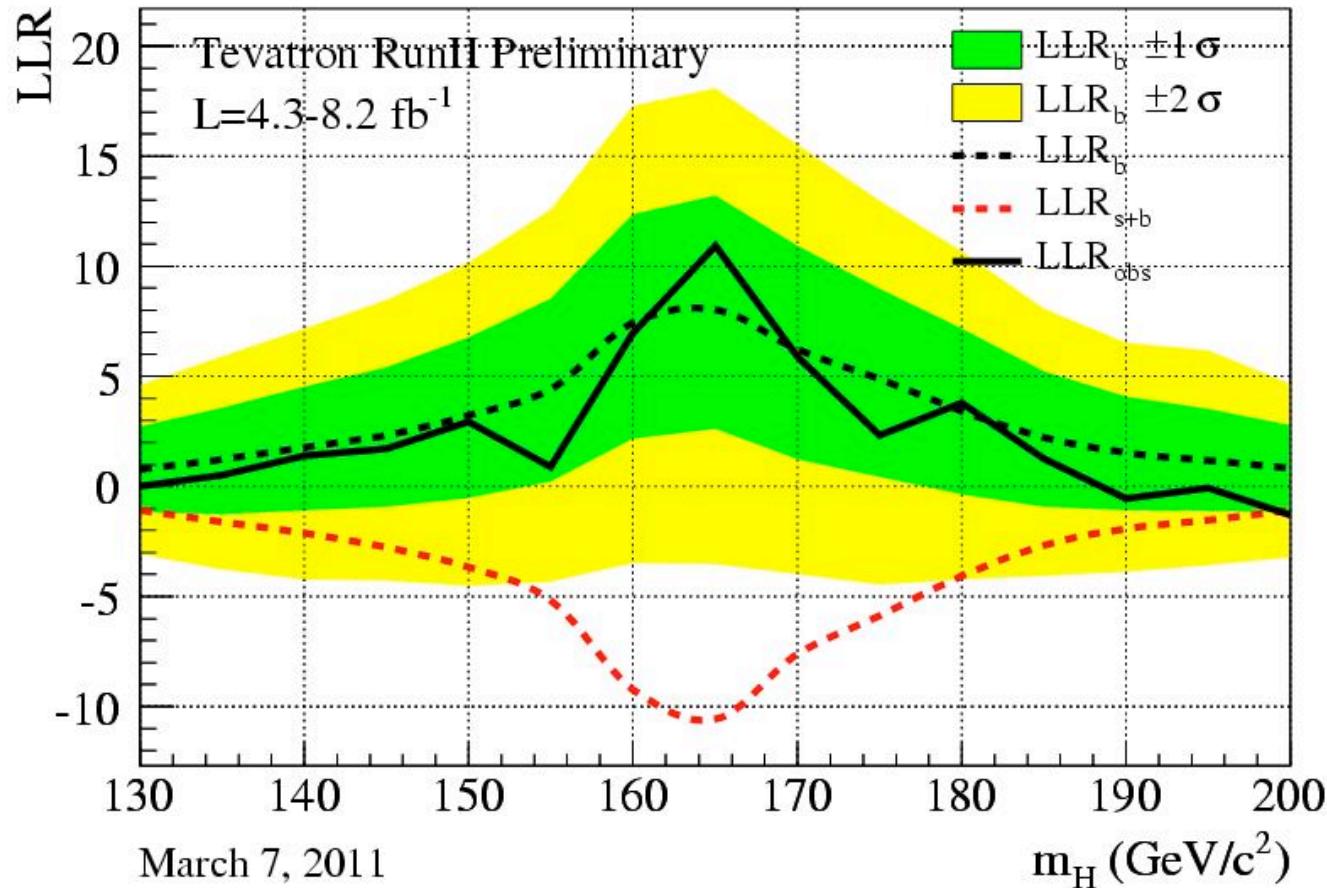
T. Junk Tevatron Higgs Combination BNL May 2011

# What These Look Like for a $5.0\sigma$ Observation



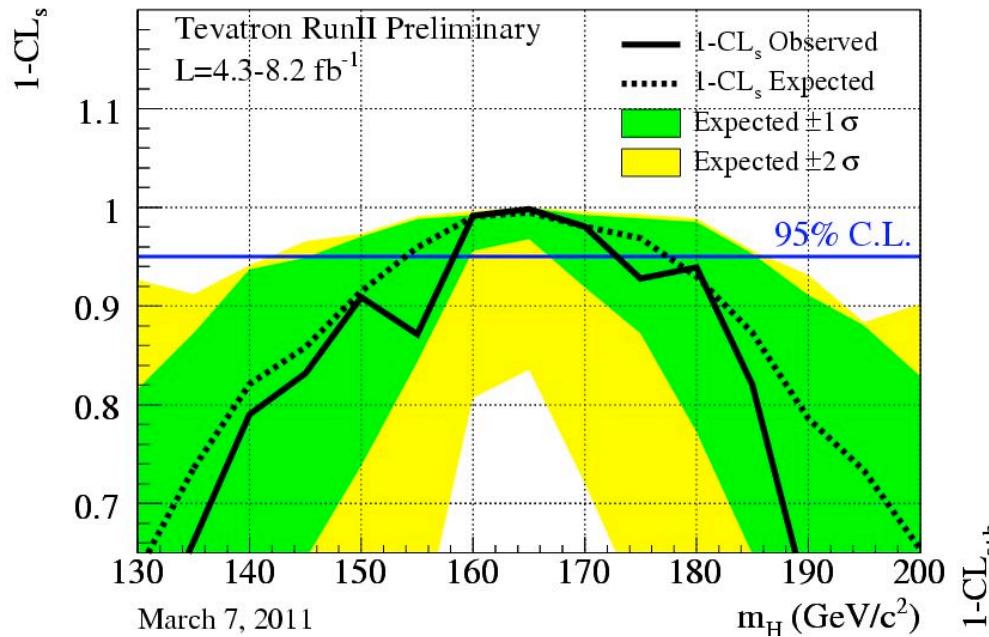
CDF Single Top,  $3.2 \text{ fb}^{-1}$

# Looking for a Signal



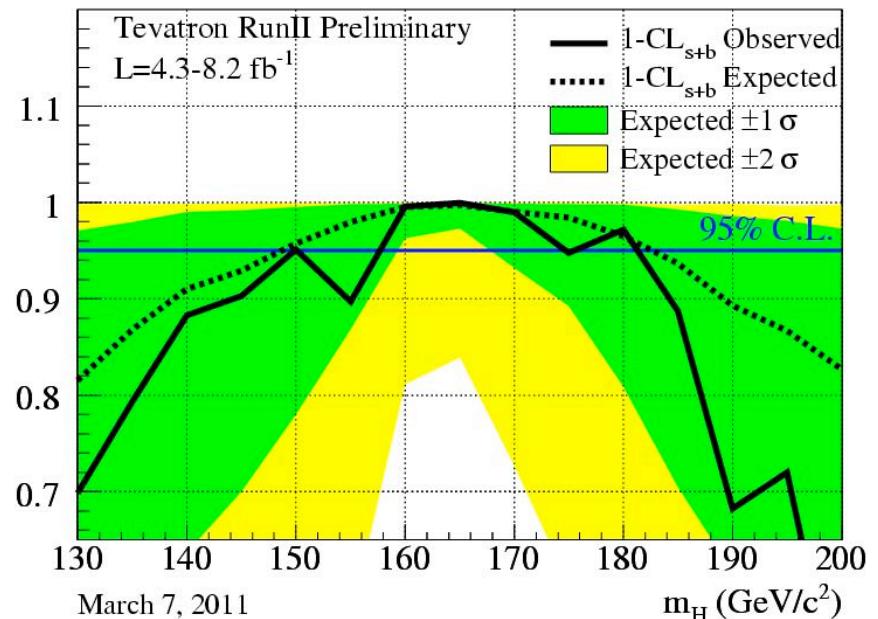
$$-2\ln Q \equiv LLR \equiv -2\ln \left( \frac{L(\text{data} | s + b, \hat{\theta})}{L(\text{data} | b, \hat{\theta})} \right)$$

# $CL_s$ vs. $CL_{s+b}$



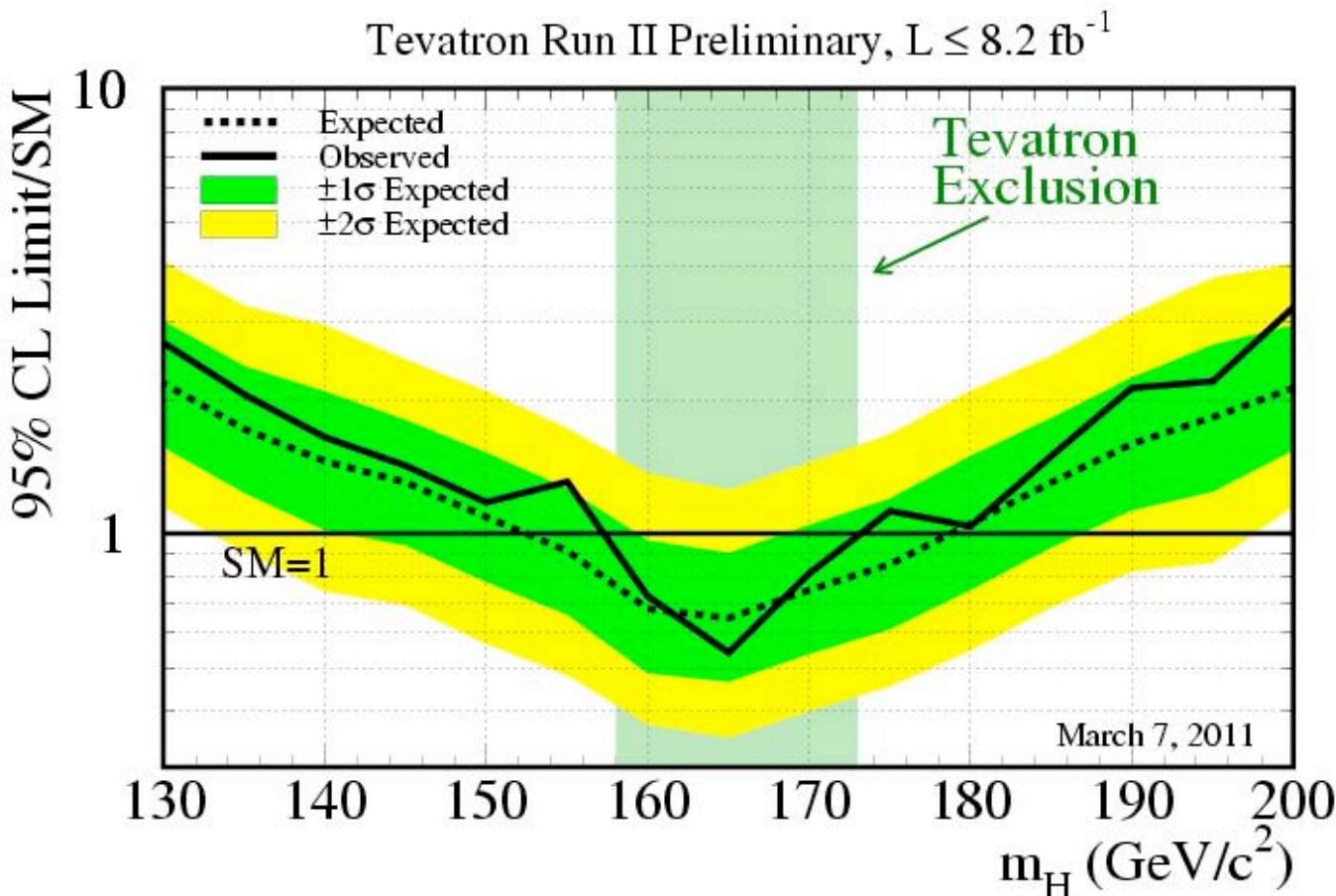
$CL_s$  matches the Bayesian limits very well

$CL_{s+b}$  and PCL give a  $\sim 40\%$  larger expected  $m_H$  exclusion



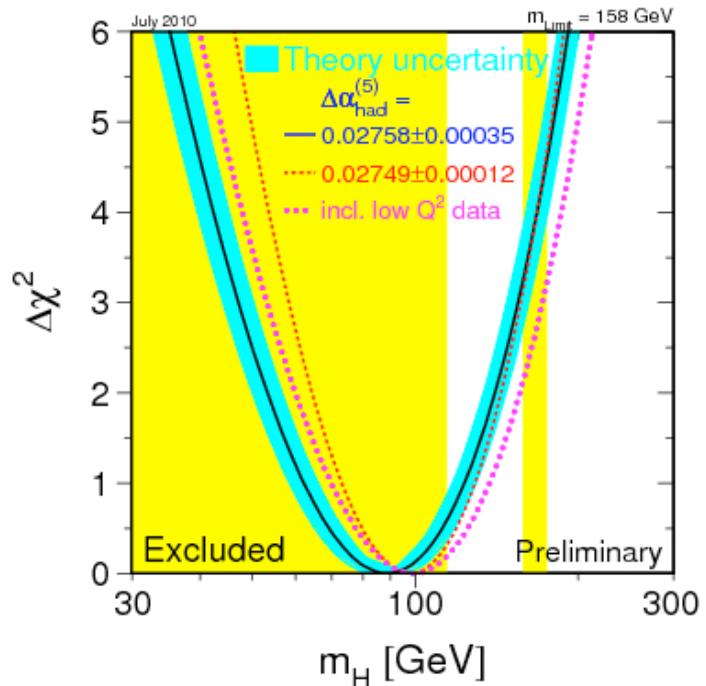
# Tevatron Observed and Expected Limits

Bayesian (chosen a priori)

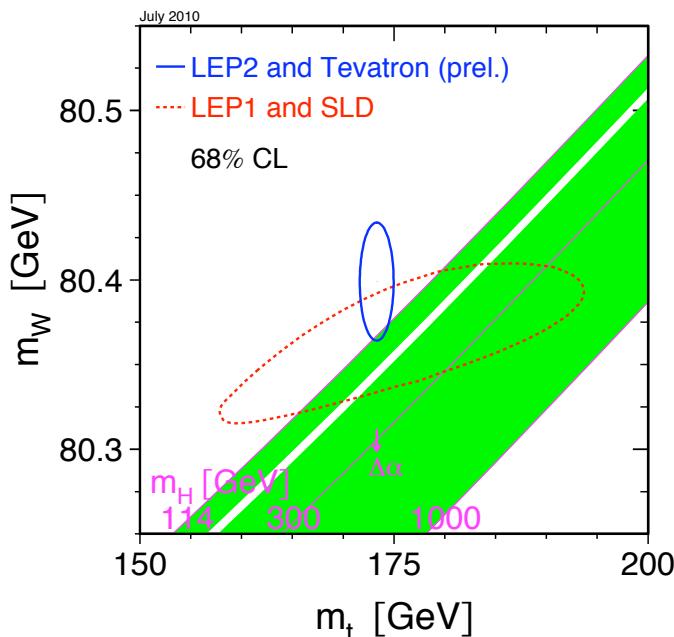


Does not include  $H \rightarrow bb$  channels – almost all  $H \rightarrow WW$  goes into this plot.  
Some  $H \rightarrow \tau\tau$  and  $H \rightarrow \gamma\gamma$  is in there however)

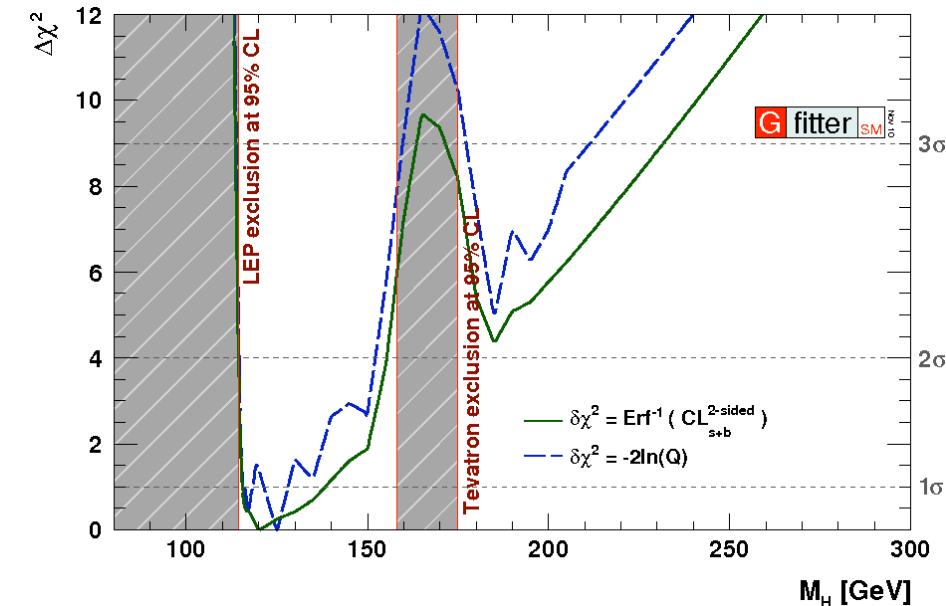
# We have a large impact on the Global Fits



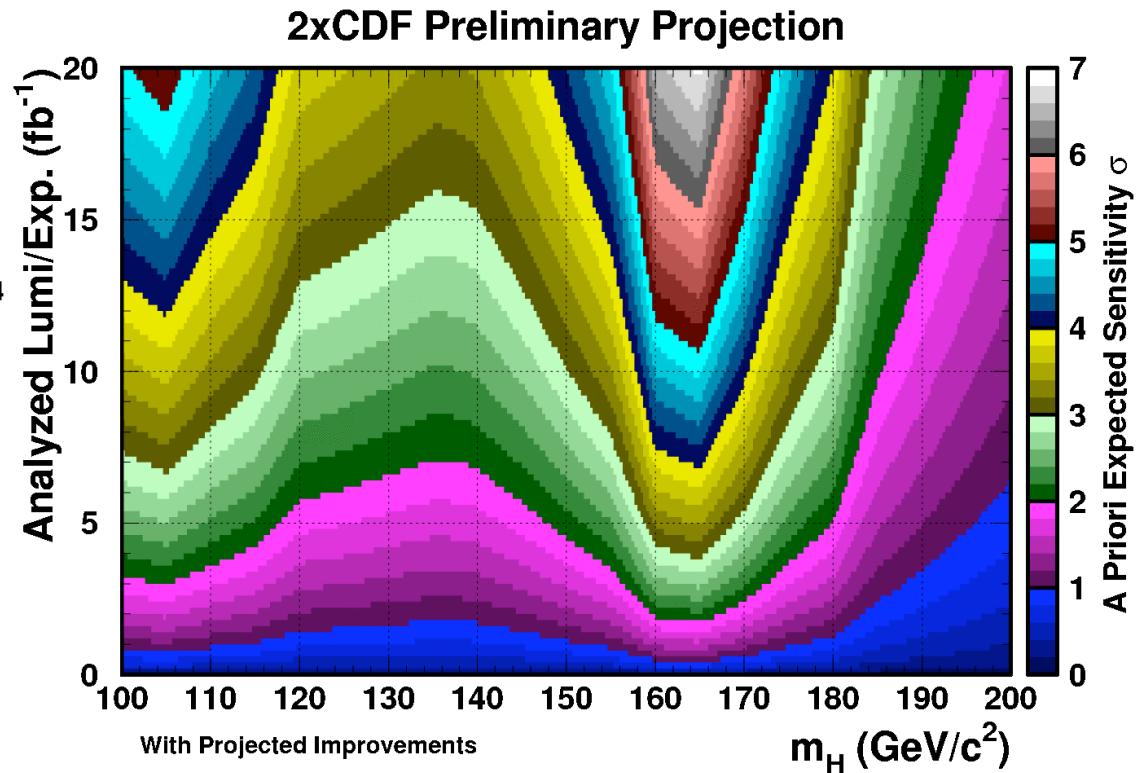
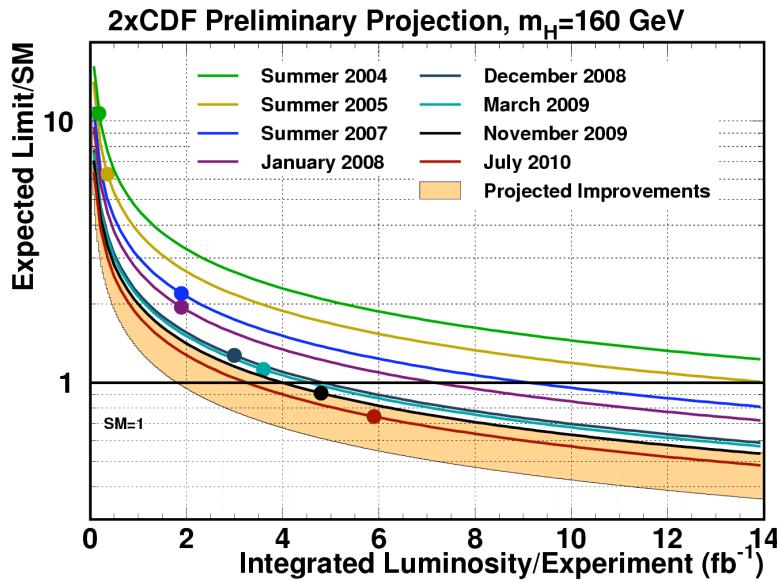
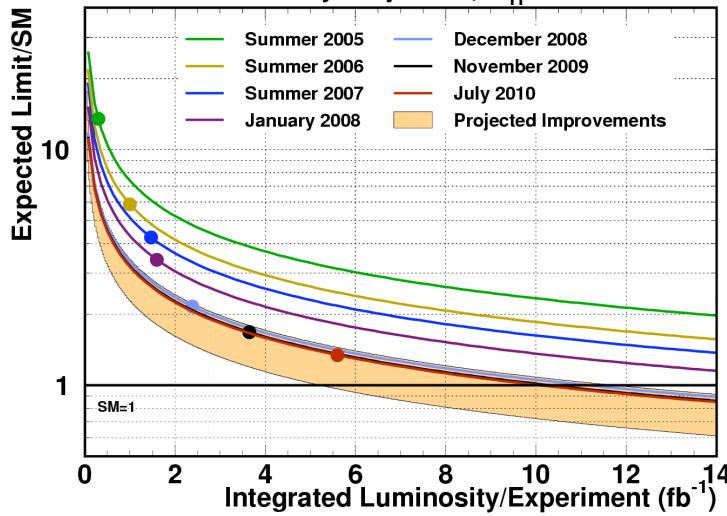
(LEPEWWG  
plots not  
updated yet)



Gfitter Collaboration:

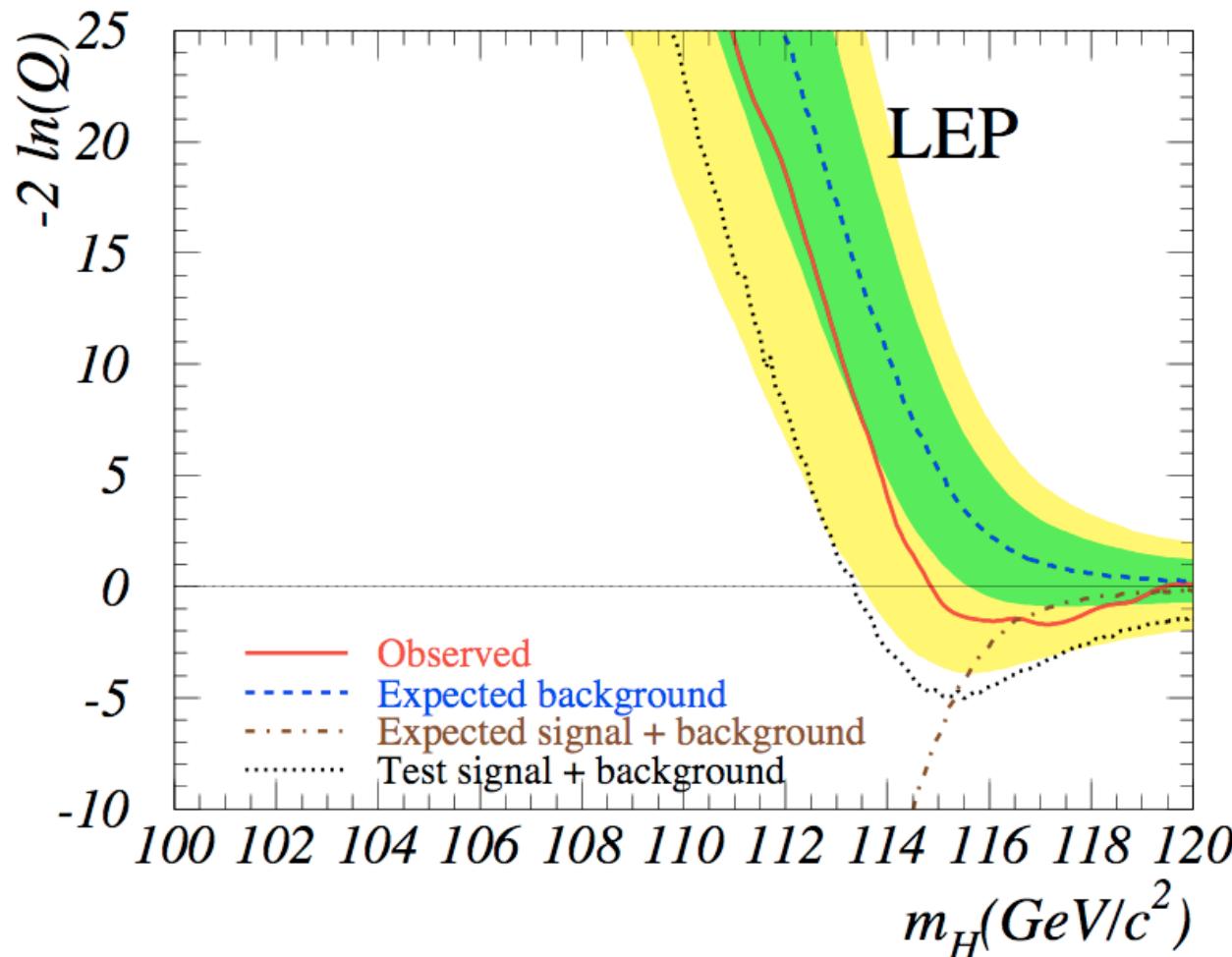


# Projections for Future Performance



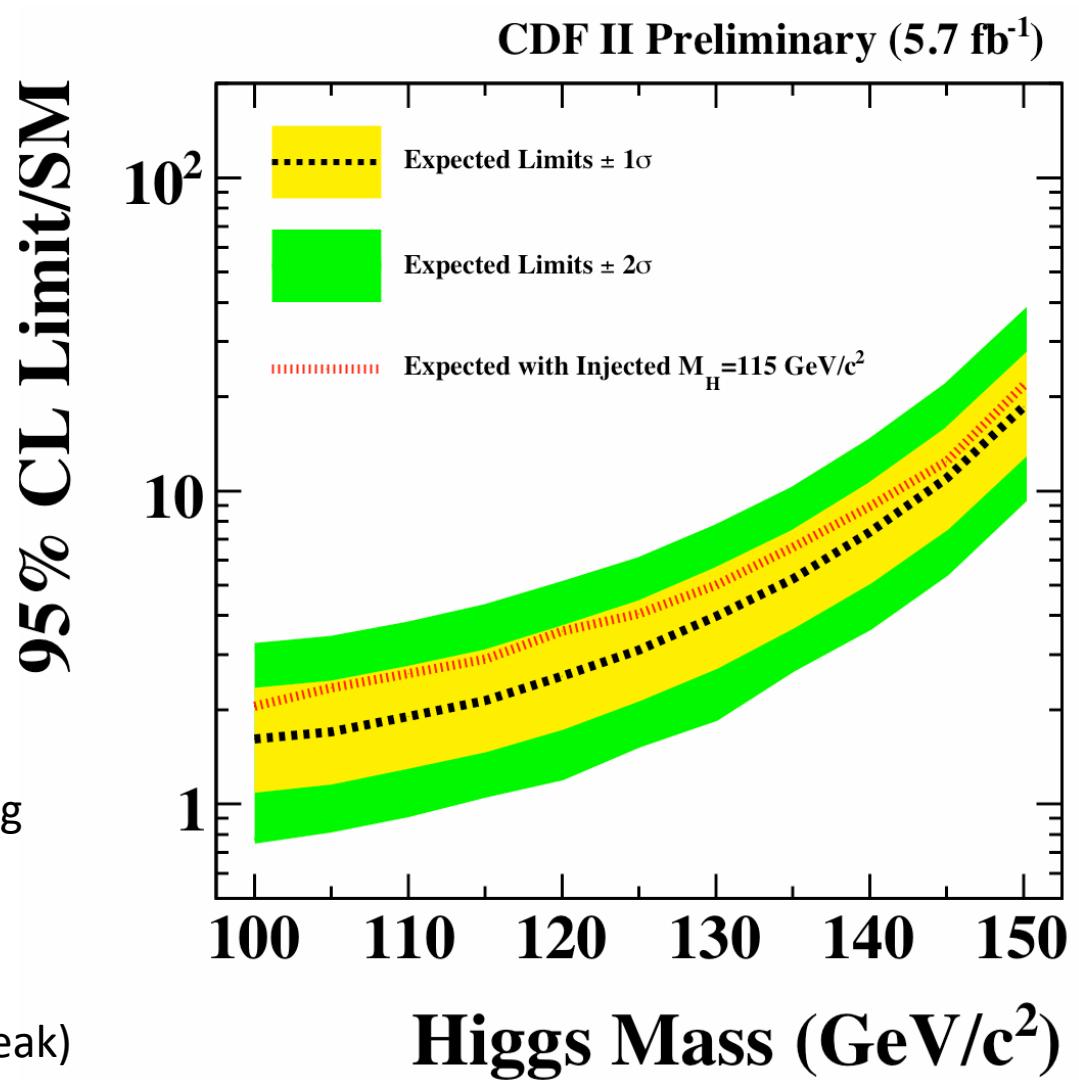
# Look-Aside Histograms

What does a signal with  $m_H = m_1$  look like when seeking  $m_H = m_2$ ?  
So far, not done at Tevatron. Not needed to study the trials factor,  
but needed to make this plot:



# Studies of Injecting a Signal at $m_H=115$ GeV

- $\ell\nu b\bar{b}$ ,  $\text{MET}b\bar{b}$ , and  $\ell\ell b\bar{b}$  channels included
- Inject SM\*1.0 signal at  $m_H=115$  GeV on top of SM backgrounds, and generate pseudoexperiments with that.
- Analyze 115 signal+background pseudoexperiments at other test masses – 100 GeV to 150 GeV
- Find the median expected limit assuming signal is there (compute it just as you would without the signal) and compare with the distribution of limits assuming the signal is completely absent.
- MVA's are less sensitive to mass than  $m_{\text{rec}}$  (fold over different sides of the peak)  
-- good for hypothesis testing, but not so good at measuring  $m_H$ .

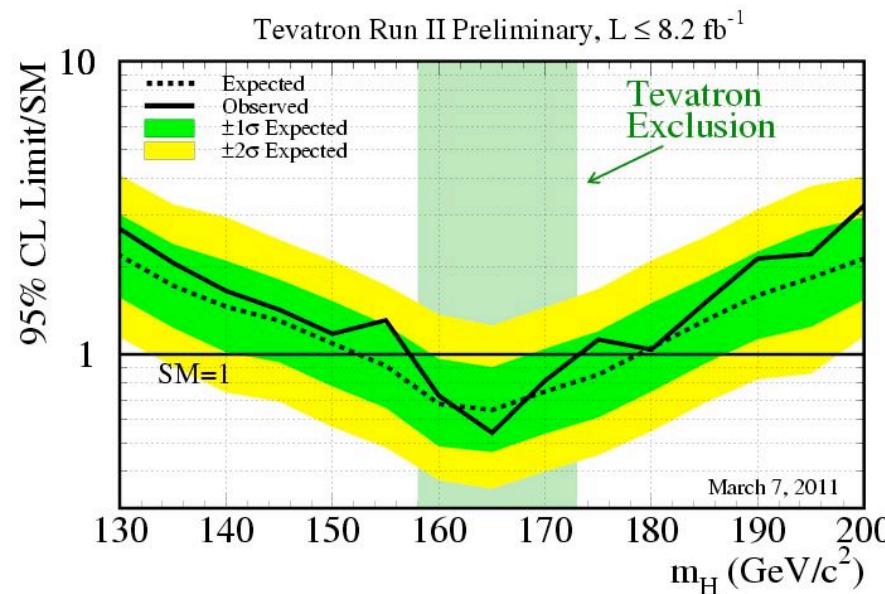


# Summary

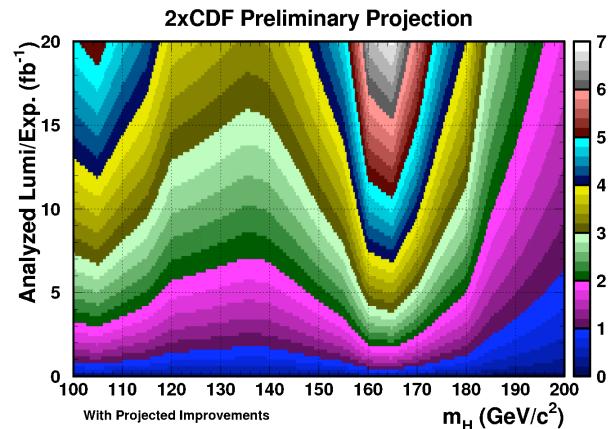
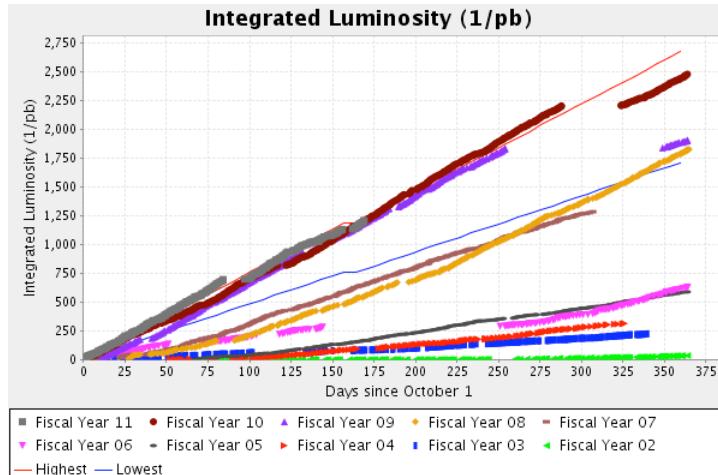
The Tevatron  
is doing very well!

We expect to  
run until the end  
of September 2011

Low- and high-mass  
updates planned for  
Summer 2011  
conferences



Short term: combine Winter 2011  $H \rightarrow \gamma\gamma$  Tevatron results



Excluded regions:

$$158 < m_H < 173 \text{ GeV}$$

Expected Exclusion  
(if no signal is present):

$$153 < m_H < 179 \text{ GeV}$$

(summer 2010  
expectation:  
 $156 < m_H < 173 \text{ GeV}$ )

# Extras

# Commonly Used Tools for Setting Limits and Discovering New Processes in use at the Tevatron

- Bayesian limits -- common at CDF
  - genlimit code by Joel Heinrich, added to mclimit code by Tom Junk. Newer Metropolis-Hastings integrator is more numerically stable on big problems
  - Implements posterior integrated over systematic uncertainties with a flat prior in cross section in 1D
  - Method described in PDG statistics review
  - Extra feature -- “correlated prior”
- $CL_s$  limits -- common at D0, but used at CDF as well.
  - Collie code by Wade Fisher in use at D0
  - Method described in PDG statistics review
  - mclimit was originally designed to do  $CL_s$  and still does.
  - TLimit in ROOT is out of date -- no fits for nuisance parameters, no shape errors or bin-by-bin errors

# Measurement and Discovery are Very Different

Buzzwords:

- Measurement = “Point Estimation”
- Discovery = “Hypothesis Testing”

You can have a discovery and a poor measurement!

Example: Expected  $b=2\times 10^{-7}$  events, expected signal=1 event, observe 1 event, no systematics.

p-value  $\sim 2\times 10^{-7}$  is a discovery! (hard to explain that event with just the background model). But have  $\pm 100\%$  uncertainty on the measured cross section!

In a one-bin search, all test statistics are equivalent. But add in a second bin, and the measured cross section becomes a poorer test statistic than the ratio of profile likelihoods.

In all practicality, discriminant distributions have a wide spectrum of s/b, even in the same histogram. But some good bins with  $b<1$  event

# Sociological Issues

- Discovery is conventionally  $5\sigma$ . In a Gaussian asymptotic case, that would correspond to a  $\pm 20\%$  measurement.
- Less precise measurements are called “measurements” all the time
- We are used to measuring undiscovered particles and processes. In the case of a background-dominated search, it can take years to climb up the sensitivity curve and get an observation, while evidence, measurements, etc. proceed.
- Referees can be confused.

# The Trials Factor

- Also called the “[Look Elsewhere Effect](#)”
- Bump-hunters are familiar with it.

What is the probability of an upward fluctuation as big as the one I saw *anywhere* in my histogram?

- Lots of bins → Lots of chances at a false discovery
- Approximation: Multiply smallest p-value by the number of “independent” models sought (not histogram bins!).

Bump hunters: roughly  $(\text{histogram width}) / (\text{mass resolution})$

Criticisms:

Adjusted p-value can now exceed unity!

What if histogram bins are empty?

What if we seek things that have been ruled out already?

It's not bins, but the number of independent hypotheses being tested that matters!

Low mass resolution at the Tevatron – not very many independent excesses possible.

# The Trials Factor

More seriously, what to do if the p-value comes from a big combination of many channels each optimized at each  $m_H$  sought?

- Channels have different resolutions (or is resolution even the right word for a multivariate discriminant?)
- Channels vary their weight in the combination as cross sections and branching ratios change with  $m_H$

Proper treatment -- want a p-value of p-values!

(use the p-value as a test statistic)

Run pseudoexperiments and analyze each one at each  $m_H$  studied. Look for the distribution of smallest p-values.

Difficult but possible.

No trials factor associated with limits. -- random oscillations tend to make excluded regions shorter the better the resolution is.

## At $m_H=165$ GeV, 133 named sources of systematic uncertainty

LUMI 0.0265679 +- 0.824895  
 D0\_Lumi 0.185264 +- 0.85186  
 D0\_MUID 0.0416735 +- 0.599  
 D0\_muontrig 0.116747 +- 0.796573  
 D0\_TauES -0.0423805 +- 0.771668  
 D0\_tautrk -0.0168425 +- 0.995416  
 D0\_TAUID 0.0078644 +- 0.941757  
 D0\_VCJ 0.36824 +- 0.550693  
 ggHpt 0.00397764 +- 0.997253  
 D0\_jetpt -0.0299649 +- 0.919487  
 D0\_JES -0.665801 +- 0.245962  
 D0\_JSSR 0.392235 +- 0.190582  
 D0\_JetID 0.294897 +- 0.278994  
 XS\_VH -0.00377098 +- 0.99713  
 XS\_qqH 0.00628008 +- 1.00377  
 XS\_ggH\_Scale -0.0678625 +- 1.00207  
 XS\_ggH\_PDF -0.0157159 +- 1.00043  
 XS\_ttbar 0.228554 +- 0.732712  
 XS\_wjets 0.214818 +- 0.98896  
 XS\_Zlp -0.949734 +- 0.750922  
 DIBOSON 0.447481 +- 0.701948  
 D0\_MJnorm -0.279252 +- 0.98592  
 D0\_MJshape 1.12234 +- 0.782589  
 D0\_pdf1 0.553438 +- 0.791879  
 XS\_Wlp 0.00432005 +- 0.610897  
 D0\_hwwmnjj2BQCD -0.678378 +- 0.620833  
 D0\_AlphaMLM -1.00197 +- 0.934272  
 D0\_AlphaScale 0.714022 +- 0.837994  
 D0\_AlphaUE -2.12227 +- 0.876342  
 D0\_hwwlnjj\_WPT -1.52096 +- 0.852375  
 D0\_hwwlnjj\_DR -0.790806 +- 0.990347  
 D0\_hwwlnjj\_JET1 0.216373 +- 0.991425  
 D0\_hwwlnjj\_JETO 0.118625 +- 0.996596  
 D0\_hwwmnjj\_TRGS 0.190183 +- 0.949918  
 D0\_TOPJES -0.00946262 +- 0.997783

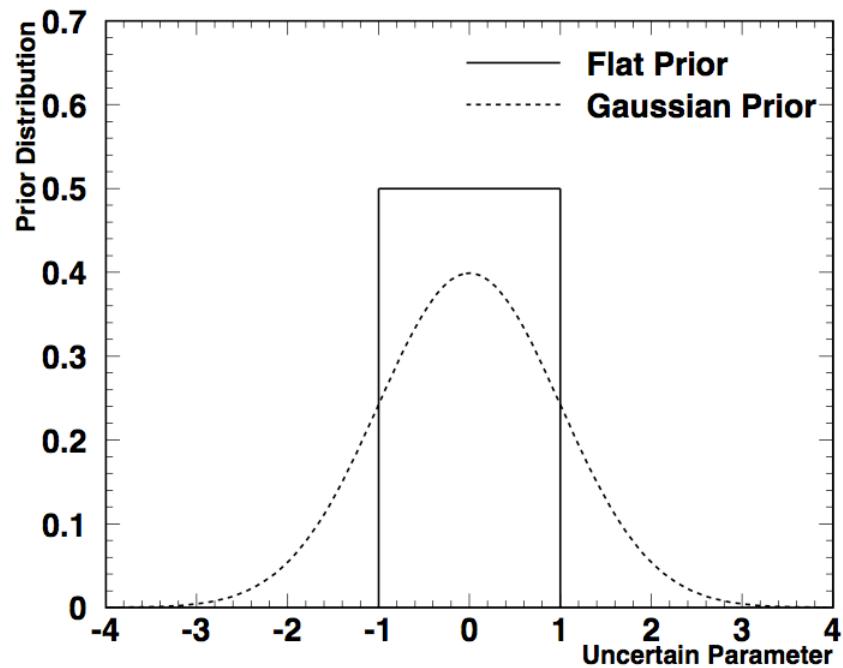
XS\_Whf -0.177845 +- 0.843819  
 XS\_Zhf -0.285996 +- 0.994315  
 D0\_BJES 0.0390238 +- 0.997961  
 D0\_hwwmnjj2AQCD -0.455895 +- 0.741364  
 D0\_EMID 0.0374239 +- 0.601478  
 D0\_hwwenjj2BQCD -1.63739 +- 0.773738  
 D0\_hwwenjj2AQCD -1.09734 +- 0.884494  
 D0\_SurfNorm\_ee2B 0.238707 +- 0.688714  
 D0\_btagOthers 1.01058 +- 0.510168  
 D0\_btagTbar -0.57316 +- 0.901125  
 D0\_gg2ww 0.0436752 +- 0.979943  
 XS\_W+lp\_jets -0.385794 +- 0.984607  
 D0\_njet2Scale2B\_wjets -0.749523 +- 0.73994  
 D0\_QCD\_hwwee2B\_2jet 0.00997179 +- 1.00411  
 D0\_njet2Scale2B\_zjets -0.205457 +- 0.422965  
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 ggHpt -0.00172292 +- 1.00173  
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 D0\_QCD\_hwwee2A\_2jet -0.0548321 +- 0.99015  
 D0\_njet2Scale2A\_zjets -0.269128 +- 0.612472  
 D0\_njet1Scale2B\_wjets -0.880726 +- 0.709565  
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 D0\_njet1Scale2B\_zjets -0.765128 +- 0.486429  
 D0\_njet1Scale2A\_wjets 0.293885 +- 0.90222  
 D0\_QCD\_hwwee2A\_1jet 0.00914123 +- 1.00193  
 D0\_njet1Scale2A\_zjets 1.26167 +- 0.728984  
 D0\_WWrew -0.0471792 +- 0.991735  
 XS\_W+lp\_Ojet -0.50853 +- 0.961063  
 D0\_njet0Scale2B\_wjets -0.0928367 +- 0.783232  
 D0\_QCD\_hwwee2B\_Ojet -0.129048 +- 0.997128  
 D0\_njet0Scale2B\_zjets -0.0250646 +- 0.910985  
 D0\_njet0Scale2A\_wjets -1.61098 +- 0.700075  
 D0\_njet0Scale2A\_zjets -0.0828907 +- 0.986453  
 D0\_SurfNorm\_mumu2B -0.469062 +- 0.625648  
 D0\_QCD\_hwwmumu -0.106863 +- 0.795989  
 D0\_Wpt -0.184552 +- 0.96306  
 D0\_Zpt 0.188088 +- 0.927037

D0\_SurfNorm\_mumu2A -0.0179862 +- 0.733734  
 D0\_MUJSSR -0.121202 +- 0.948451  
 D0\_etrig -0.0351927 +- 0.985838  
 D0\_WH\_JET2B -0.00347408 +- 0.99674  
 D0\_WH\_WJET2B -0.0523859 +- 0.958982  
 D0\_WHME\_QCD2B -0.112805 +- 0.934644  
 D0\_WHMM\_QCD2B 0.00983068 +- 0.988268  
 D0\_WHMM\_FLIP2B -0.228278 +- 0.747162  
 D0\_WHEE\_QCD2B 0.0364418 +- 0.984796  
 D0\_WHEE\_FLIP2B 0.0224206 +- 1.00059  
 D0\_WH\_JET2A -0.020881 +- 0.996993  
 D0\_WH\_WJET2A -0.111409 +- 1.00083  
 D0\_WHME\_QCD2A 0.414989 +- 0.912829  
 D0\_WHMM\_QCD2A 0.0119617 +- 0.938078  
 D0\_WHMM\_FLIP2A -0.00789962 +- 0.650085  
 D0\_WHEE\_QCD2A -0.225438 +- 0.940028  
 D0\_WHEE\_FLIP2A 0.0202663 +- 0.930071  
 D0\_HWWemu\_trigger -0.0584725 +- 0.878939  
 D0\_SurfNorm\_em2B -0.0717138 +- 0.816335  
 D0\_QCD\_hwwmem2B\_2jet -0.360772 +- 0.97699  
 D0\_EMJSSR -0.00838239 +- 1.00116  
 D0\_QCD\_hwwmem2B\_1jet 0.10249 +- 0.982492  
 D0\_WJETS\_SHAPE\_ss2B\_Ojet 1.36122 +- 0.50868  
 D0\_QCD\_hwwmem2B\_Ojet -1.00002 +- 0.897224  
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 D0\_QCD\_hwwmemu 0.110679 +- 0.986306  
 D0\_njet2Scale 0.104044 +- 0.960317  
 D0\_njet1Scale 0.135863 +- 0.777831  
 D0\_njet0Scale 0.608654 +- 0.685043  
 D0\_Triggers -0.00535372 +- 1.00348  
 D0\_Wjet\_mutau 0.298426 +- 0.490672  
 D0\_QCD\_mutau 0.518902 +- 0.622592  
 TrigID 0.020004 +- 0.895659  
 WWptScale -0.196794 +- 0.981869  
 WWptPDF 0.00458475 +- 0.961389  
 CDFLUMI 0.0382205 +- 0.866633

## At $m_H=165$ GeV, 133 named sources of systematic uncertainty

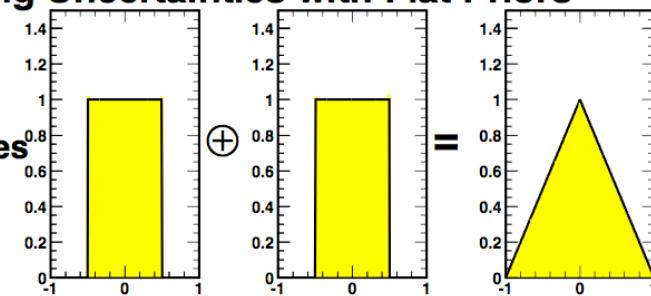
CDFLEPTONID 0.000794111 +- 0.994515  
CDFJES 0.370398 +- 0.37521  
CDFHWW\_TopAcc 0.731538 +- 0.806684  
CDFHWW\_DiboAcc 1.16657 +- 0.906918  
WgamScale 0.792679 +- 0.76395  
CDFHWW\_WgamAcc 0.359372 +- 0.92247  
CDFHWW\_FakeRate -0.0250224 +- 0.475505  
METModel -0.951287 +- 0.388791  
HetaPDF -0.0155761 +- 0.996384  
HptPDF 0.00182511 +- 0.994836  
ALPHAS 0.000351948 +- 1.0022  
MB 5.43305e-05 +- 1.00693  
MC 0.00971472 +- 1.00016  
ttbar\_bfake 0.0156202 +- 0.971993  
XS\_Vgamma -0.0833847 +- 0.995168  
ZgamAcc -0.114853 +- 0.973599  
TauJetFake 0.189097 +- 0.83341  
XS\_Wjet -0.1286 +- 0.678103  
TauLepFake 0.0505157 +- 1.0017  
XS\_DY -0.0257132 +- 0.99126  
CDFHWW\_DYAcc -0.0609454 +- 0.950461  
TauID 0.12933 +- 0.998585  
WW\_CFR -0.277979 +- 0.911791  
CDFHWW\_BTagVeto 0.0775033 +- 0.972641

Constraints checked twice  
1) MINUIT output  
2) Bayesian posterior (by-product of Markov Chain integration)

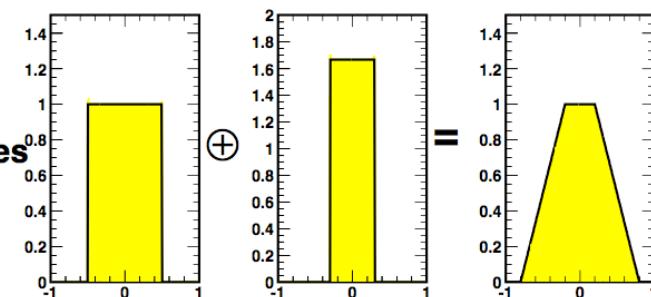


### Adding Uncertainties with Flat Priors

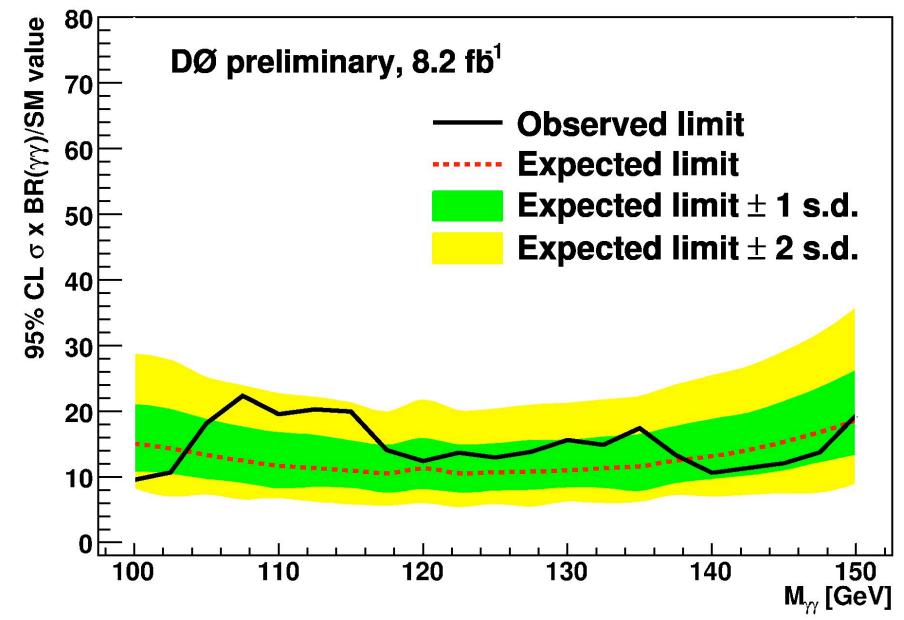
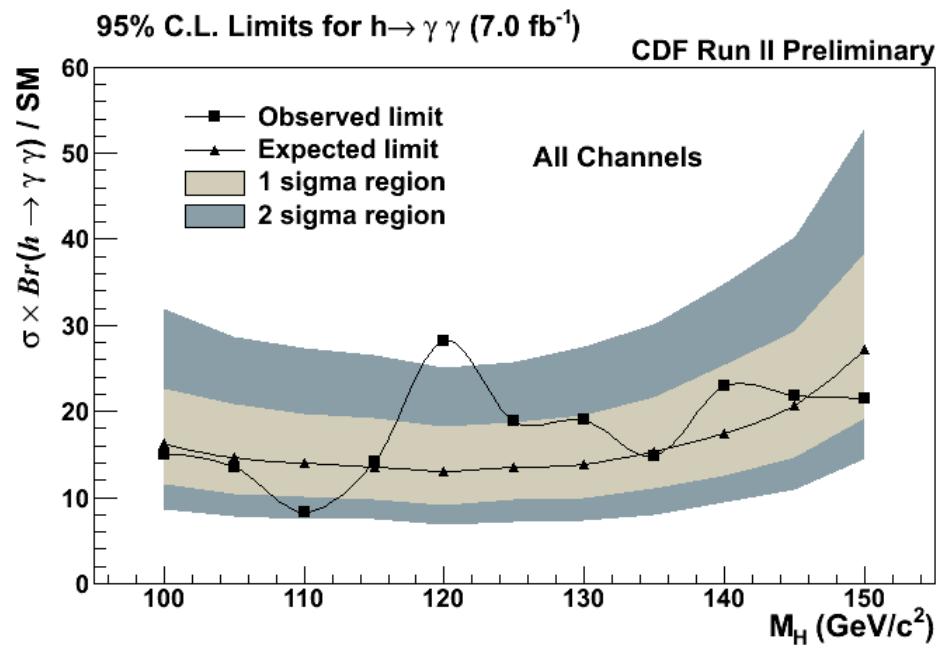
Equal  
Uncertainties



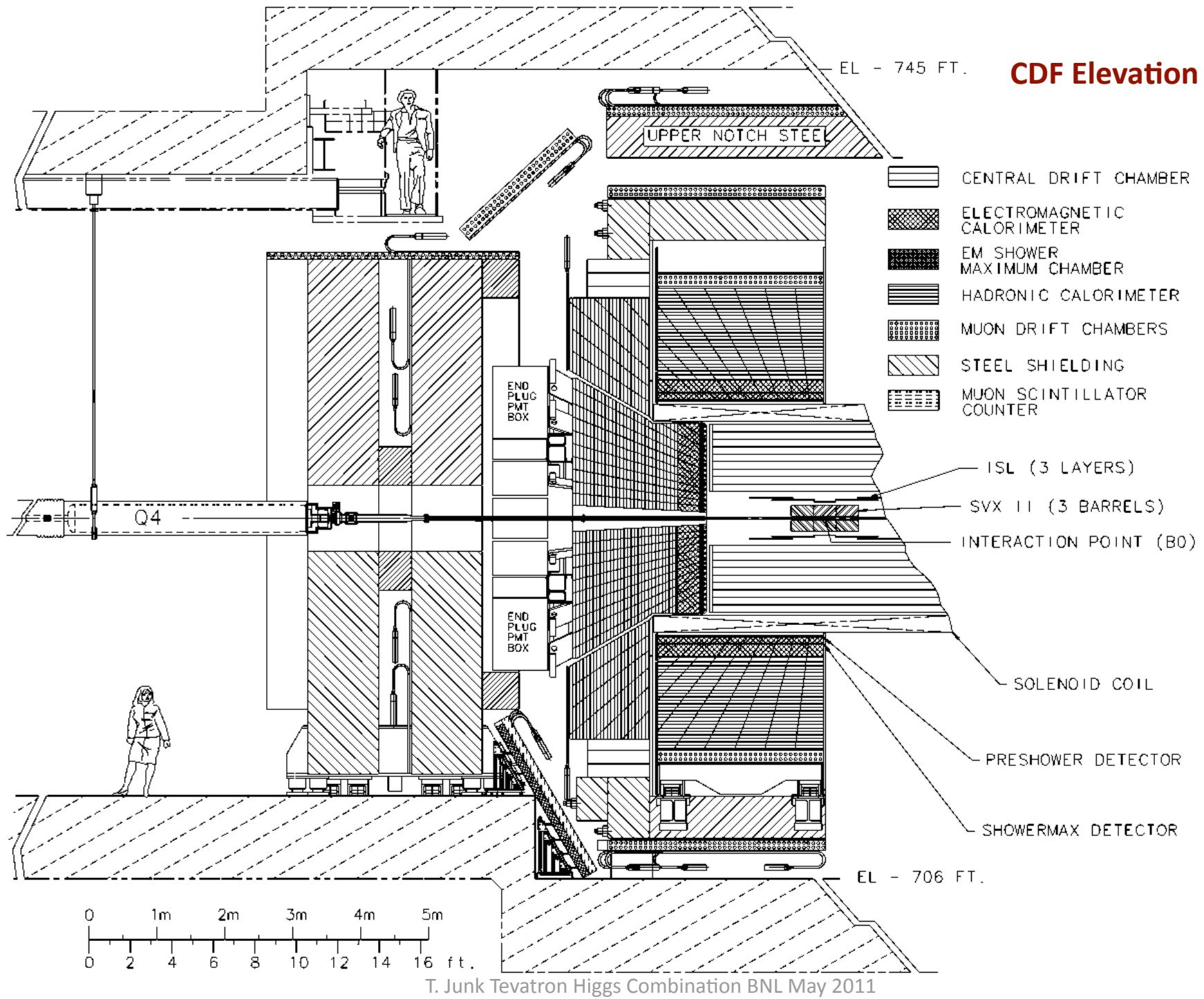
Unequal  
Uncertainties



# CDF and D0 $H \rightarrow \gamma\gamma$ Limits



## CDF Elevation



## Integrating the Posterior to Other Values than 95%

